

New Opportunities in Neutrino Physics

University of Colorado - Seminar

Milind Diwan 11/9/2004

Brookhaven National Laboratory

Upton NY 11776

- Very Brief Review
- Description of 3-generation oscillations
- What to expect in 5 years
- Ambitions for DUSEL !

Key Items: It is possible to build an intense beam and it makes sense to send it more than 2000 km

Thanks to many for slides. esp: SK,
SNO, Kamland, Minos, and APS
neutrino study



Standard Model of FUNDAMENTAL PARTICLES AND INTERACTIONS

The Standard Model summarizes the current knowledge in Particle Physics. It is the quantum theory that includes the theory of strong interactions (quantum chromodynamics or QCD) and the unified theory of weak and electromagnetic interactions (electroweak). Gravity is included on this chart because it is one of the fundamental interactions even though not part of the "Standard Model."

FERMIONS

matter constituents
spin = 1/2, 3/2, 5/2, ...

Leptons spin = 1/2			Quarks spin = 1/2		
Flavor	Mass GeV/c ²	Electric charge	Flavor	Approx. Mass GeV/c ²	Electric charge
ν_e electron neutrino	$<1 \times 10^{-8}$	0	u up	0.003	2/3
e electron	0.000511	-1	d down	0.006	-1/3
ν_μ muon neutrino	<0.0002	0	c charm	1.3	2/3
μ muon	0.106	-1	s strange	0.1	-1/3
ν_τ tau neutrino	<0.02	0	t top	175	2/3
τ tau	1.7771	-1	b bottom	4.3	-1/3

Spin is the intrinsic angular momentum of particles. Spin is given in units of \hbar , which is the quantum unit of angular momentum, where $\hbar = h/2\pi = 6.58 \times 10^{-25}$ GeV s = 1.05×10^{-34} J s.

Electric charges are given in units of the proton's charge. In SI units the electric charge of the proton is 1.60×10^{-19} coulombs.

The **energy** unit of particle physics is the electronvolt (eV), the energy gained by one electron in crossing a potential difference of one volt. **Masses** are given in GeV/c² (remember $E = mc^2$), where 1 GeV = 10^9 eV = 1.60×10^{-10} joule. The mass of the proton is 0.938 GeV/c² = 1.67×10^{-27} kg.

BOSONS

force carriers
spin = 0, 1, 2, ...

Unified Electroweak spin = 1			Strong (color) spin = 1		
Name	Mass GeV/c ²	Electric charge	Name	Mass GeV/c ²	Electric charge
γ photon	0	0	g gluon	0	0
W^-	80.4	-1			
W^+	80.4	+1			
Z^0	91.187	0			

Color Charge

Each quark carries one of three types of "strong charge," also called "color charge." These charges have nothing to do with the colors of visible light. There are eight possible types of color charge for gluons. Just as electrically charged particles interact by exchanging photons, in strong interactions color-charged particles interact by exchanging gluons. Leptons, photons, and W and Z bosons have no strong interactions and hence no color charge.

Quarks Confined in Mesons and Baryons

One cannot isolate quarks and gluons; they are confined in color-neutral particles called **hadrons**. This confinement (binding) results from multiple exchanges of gluons among the color-charged constituents. As color-charged particles (quarks and gluons) move apart, the energy in the color-force field between them increases. This energy eventually is converted into additional quark-antiquark pairs (see figure below). The quarks and antiquarks then combine into hadrons; these are the particles seen to emerge. Two types of hadrons have been observed in nature: **mesons** $q\bar{q}$ and **baryons** qqq .

Residual Strong Interaction

The strong binding of color-neutral protons and neutrons to form nuclei is due to residual strong interactions between their color-charged constituents. It is similar to the residual electrical interaction that binds electrically neutral atoms to form molecules. It can also be viewed as the exchange of mesons between the hadrons.

PROPERTIES OF THE INTERACTIONS

Baryons qqq and Antibaryons $\bar{q}\bar{q}\bar{q}$					
Baryons are fermionic hadrons. There are about 120 types of baryons.					
Symbol	Name	Quark content	Electric charge	Mass GeV/c ²	Spin
p	proton	uud	1	0.938	1/2
\bar{p}	anti-proton	$\bar{u}\bar{u}\bar{d}$	-1	0.938	1/2
n	neutron	udd	0	0.940	1/2
Λ	lambda	uds	0	1.116	1/2
Ω^-	omega	sss	-1	1.672	3/2

Property \ Interaction	Gravitational	Weak	Electromagnetic	Strong	
		(Electroweak)		Fundamental	Residual
Acts on:	Mass – Energy	Flavor	Electric Charge	Color Charge	See Residual Strong Interaction Note
Particles experiencing:	All	Quarks, Leptons	Electrically charged	Quarks, Gluons	Hadrons
Particles mediating:	Graviton (not yet observed)	W^+ W^- Z^0	γ	Gluons	Mesons
Strength relative to electromag for two u quarks at:	10^{-41}	0.8	1	25	Not applicable to quarks
	10^{-41}	10^{-4}	1	60	
	10^{-36}	10^{-7}	1	Not applicable to hadrons	
for two protons in nucleus				20	

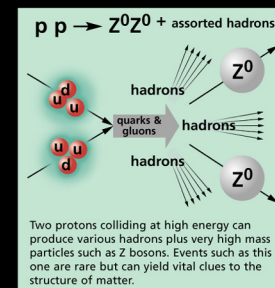
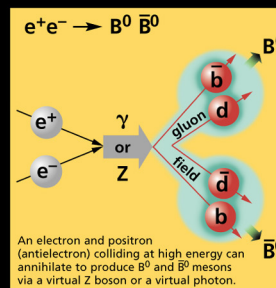
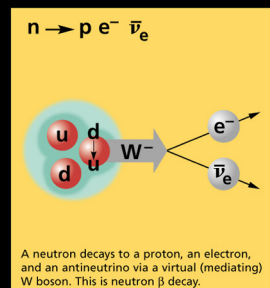
Mesons $q\bar{q}$					
Mesons are bosonic hadrons. There are about 140 types of mesons.					
Symbol	Name	Quark content	Electric charge	Mass GeV/c ²	Spin
π^+	pion	u \bar{d}	+1	0.140	0
K^-	kaon	s \bar{u}	-1	0.494	0
ρ^+	rho	u \bar{d}	+1	0.770	1
B^0	B-zero	d \bar{b}	0	5.279	0
η_c	eta-c	c \bar{c}	0	2.980	0

Matter and Antimatter

For every particle type there is a corresponding antiparticle type, denoted by a bar over the particle symbol (unless + or - charge is shown). Particle and antiparticle have identical mass and spin but opposite charges. Some electrically neutral bosons (e.g., Z^0 , γ , and η_c = $c\bar{c}$, but not K^0 = $d\bar{s}$) are their own antiparticles.

Figures

These diagrams are an artist's conception of physical processes. They are **not** exact and have **no** meaningful scale. Green shaded areas represent the cloud of gluons or the gluon field, and red lines the quark paths.



The Particle Adventure

Visit the award-winning web feature *The Particle Adventure* at <http://ParticleAdventure.org>

This chart has been made possible by the generous support of:

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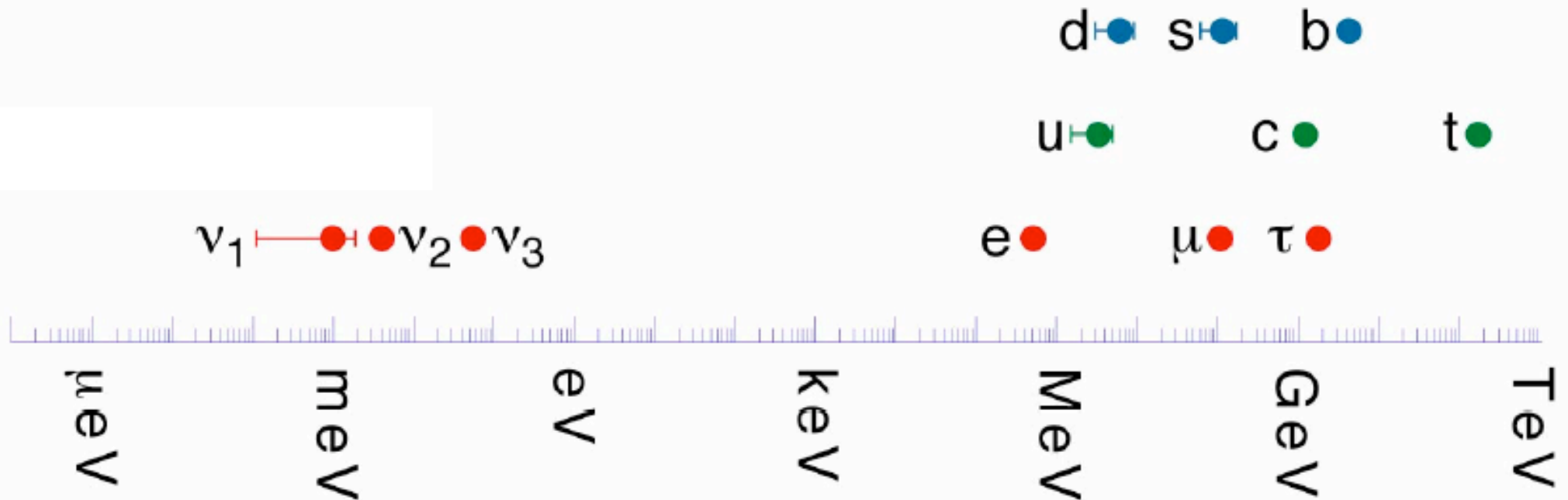
<http://CPEPweb.org>

Neutrino puzzles

- Do they have mass ? Why so small ?
- If they have mass what implications on left-right properties ?
- Can they turn into each other ?
- What implications for the structure of the universe ?
- What is the relationship to quarks ?

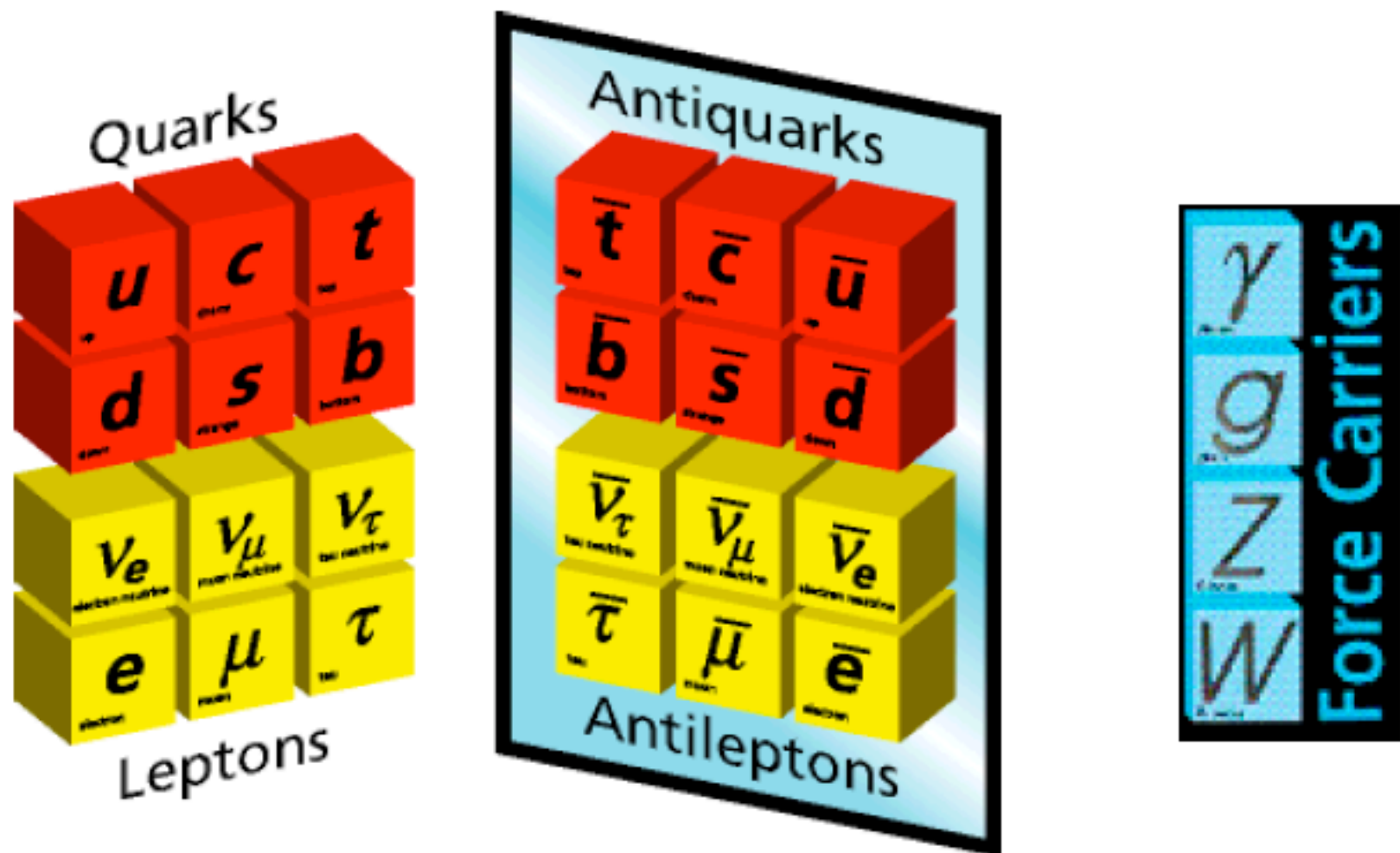
Current picture of masses from oscillations puzzling.

fermion masses



hierarchy

The Standard Model



This picture needs revision

Why Mass could imply Lepton number violation

	Particle	Anti-particle
Left	$(e \quad \nu)_L$	$\overline{(e \quad \nu)_L}$
Right	$e_R \quad \nu_R$	$\bar{e}_R \quad \bar{\nu}_R$

- Standard model has only left handed leptons in isospin states. But if neutrino has mass it can become right handed.
- If $\bar{\nu}_L = \nu_R$ then neutrinos are their own antiparticles and can annihilate themselves.

Brief review of oscillations

Assume a 2×2 neutrino mixing matrix.

$$\begin{pmatrix} \nu_a \\ \nu_b \end{pmatrix} = \begin{pmatrix} \cos(\theta) & \sin(\theta) \\ -\sin(\theta) & \cos(\theta) \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

$$\nu_a(t) = \cos(\theta)\nu_1(t) + \sin(\theta)\nu_2(t)$$

$$\begin{aligned} P(\nu_a \rightarrow \nu_b) &= | \langle \nu_b | \nu_a(t) \rangle |^2 \\ &= \sin^2(\theta) \cos^2(\theta) |e^{-iE_2 t} - e^{-iE_1 t}|^2 \end{aligned}$$

Sufficient to understand most of the physics:

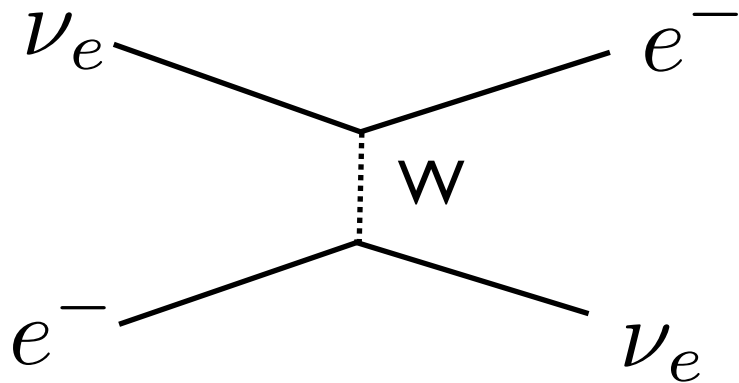
$$P(\nu_a \rightarrow \nu_b) = \sin^2 2\theta \sin^2 \frac{1.27((m_2^2 - m_1^2)/eV^2)(L/km)}{(E/GeV)}$$

$$P(\nu_a \rightarrow \nu_a) = 1 - \sin^2 2\theta \sin^2 \frac{1.27(\Delta m^2/eV^2)(L/km)}{(E/GeV)}$$

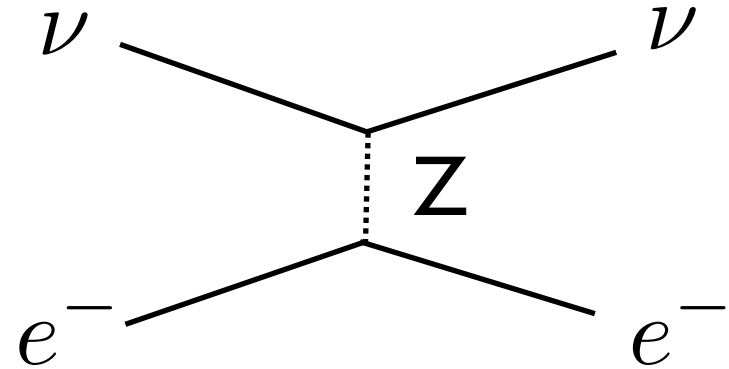
Oscillation nodes at $\pi/2, 3\pi/2, 5\pi/2, \dots$ ($\pi/2$): $\Delta m^2 = 0.0025 eV^2$,
 $E = 1 GeV$, $L = 494 km$.

$$i \frac{d}{dx} \nu_f = H R_\theta \nu_m$$

L. Wolfenstein: Oscillations need to be modified in presence of matter.



Charged Current
for electron type only



Neutral Current
for all neutrino types

Additional potential for ν_e ($\bar{\nu}_e$): $\pm \sqrt{2} G_F N_e$

N_e is electron number density.

Oscillations in presence of matter

$$i \frac{d}{dx} \nu_f = R_\theta H(\nu_m) + H_{mat}(\nu_f)$$

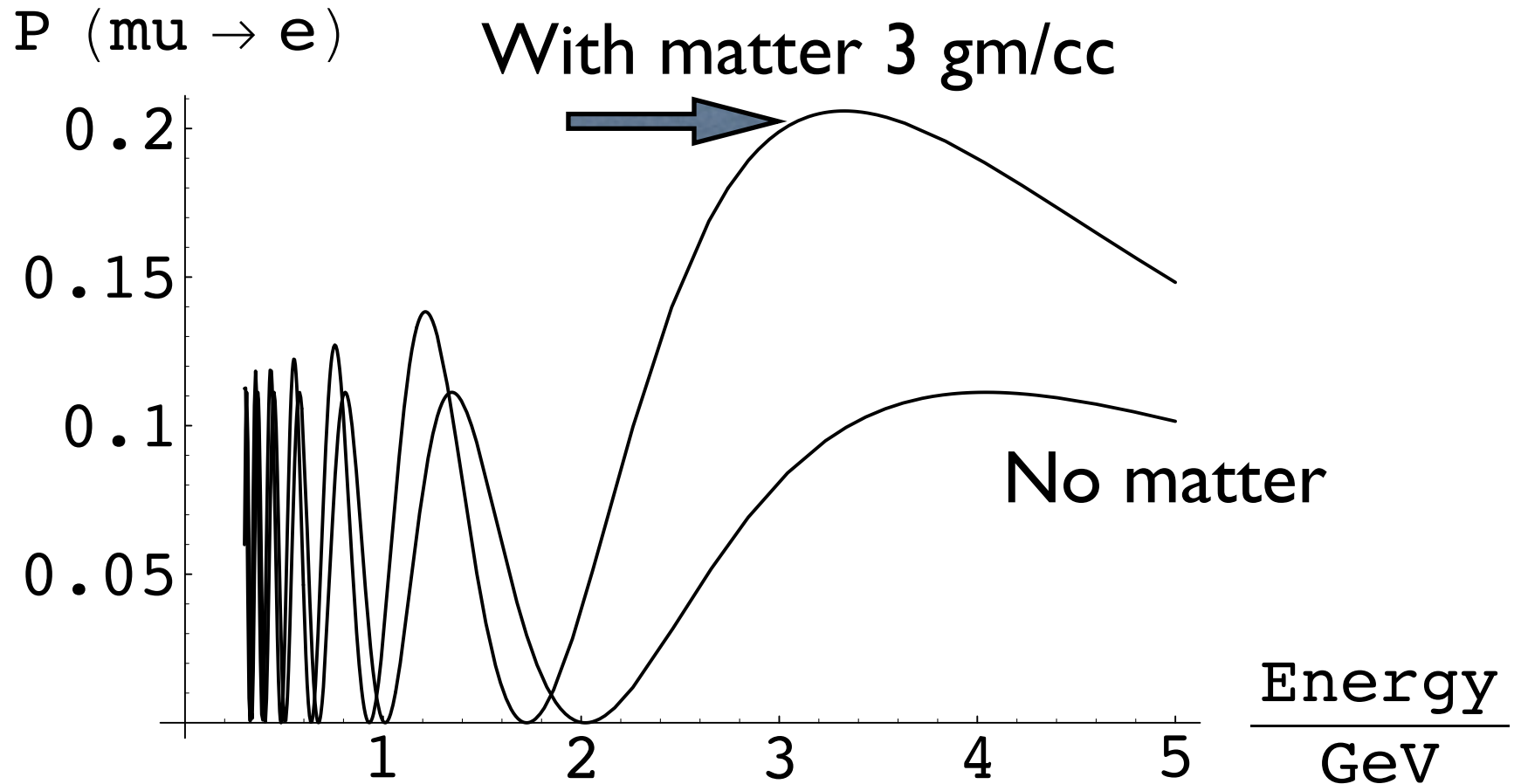
$$i \frac{d}{dx} \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \frac{1}{4E} \left(R_\theta \begin{pmatrix} m_1^2 & 0 \\ 0 & m_2^2 \end{pmatrix} R_\theta^T + 2E \begin{pmatrix} \sqrt{2}G_F N_e & 0 \\ 0 & -\sqrt{2}G_F N_e \end{pmatrix} \right) \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} \quad (3)$$

$$P_{\mu \rightarrow e} = \frac{\sin^2 2\theta}{(\cos 2\theta - a)^2 + \sin^2 2\theta} \times \sin^2 \frac{L\Delta m^2}{4E} \sqrt{(a - \cos 2\theta)^2 + \sin^2 2\theta}$$

$$\begin{aligned} a &= 2\sqrt{2}EG_F N_e / \Delta m^2 \\ &\approx 7.6 \times 10^{-5} \times D / (gm/cc) \times E_\nu / GeV / (\Delta m^2 / eV^2) \end{aligned} \quad (4)$$

Important only if electron neutrinos in the mix

2-neutrino picture



Osc. probability: 0.0025 eV^2 , $L = 2000 \text{ km}$, $\Theta = 10^\circ$

Key new evidence

- Super KamiokaNDE (SK): observe atmospheric neutrinos.
- Sudbury Neutrino Observatory (SNO): observed solar neutrinos.
- KEK to SK accelerator beam
- KAMLAND reactor experiment
- If LSND is confirmed the main picture does not change, but you need more.

Apologies to many other pioneering experiments

What do we know and how do we know it

Not known
Has CP phase

Bounded by CHOOZ

{ From Max. Atm. mixing,
 $\nu_3 \equiv (\nu_\mu + \nu_\tau) / \sqrt{2}$

Don't know sign

Ignoring
LSND

(mass)²

Δm_{atm}^2

0.0025 eV²

{ From ν_μ (Up) oscillate
but ν_μ (Down) don't

{ In LMA-MSW, $P_\odot(\nu_e \rightarrow \nu_e)$
= ν_e fraction of ν_2 and KamLAND

ν_2

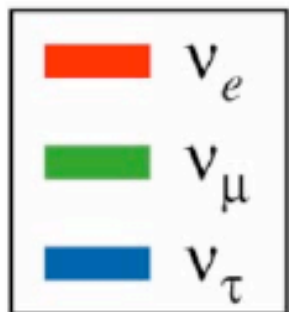
ν_1

{ From distortion of ν_e (solar)
and $\bar{\nu}_e$ (reactor) spectra

0.000008 eV²

{ From Max. Atm. mixing, ν_1 & ν_2
include $(\nu_\mu - \nu_\tau) / \sqrt{2}$

Measurements
not yet precise



Slide adapted from B. Kayser

Big picture physics issues for neutrinos

- Precision measurements

$$\Delta m_{21}^2, \sin^2 2\theta_{12}, \Delta m_{32}^2, \sin^2 2\theta_{23}$$

- Implications of 3-generation mixing.

$$\nu_\mu \rightarrow \nu_e, \sin^2 2\theta_{13}, \delta_{CP}$$

- New physics: deviations from $\sin^2 \frac{\Delta m^2 L}{4E}$,
new interactions, new symmetries. Sterile
neutrinos: (LSND-miniboone exp.)

If LSND is confirmed these goals don't change.

3-generation oscillations

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \quad (3)$$

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \quad (4)$$

3-generation formula without matter effect:

$$\begin{aligned} P(\nu_a \rightarrow \nu_b) = & \sum_i |U_{ai}|^2 |U_{bi}|^2 \\ & 2\text{Re}(U_{a1}^* U_{b1} U_{a2} U_{b2}^* \times \exp(-i\Delta m_{21}^2 L/2E) \\ & 2\text{Re}(U_{a1}^* U_{b1} U_{a3} U_{b3}^* \times \exp(-i\Delta m_{31}^2 L/2E) \\ & 2\text{Re}(U_{a2}^* U_{b2} U_{a3} U_{b3}^* \times \exp(-i\Delta m_{32}^2 L/2E) \end{aligned}$$

For anti-neutrinos take complex conjugate of matrix. Difference from 2 generations: phases.

$\nu_\mu \rightarrow \nu_e$ with matter effect

Approximate formula (Lindener, Huber et al.)

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) \approx & \sin^2 \theta_{23} \frac{\sin^2 2\theta_{13}}{(\hat{A} - 1)^2} \sin^2((\hat{A} - 1)\Delta) \\
 & + \alpha \frac{8J_{CP}}{\hat{A}(1 - \hat{A})} \sin(\Delta) \sin(\hat{A}\Delta) \sin((1 - \hat{A})\Delta) \\
 & + \alpha \frac{8I_{CP}}{\hat{A}(1 - \hat{A})} \sin(\Delta) \cos(\hat{A}\Delta) \sin((1 - \hat{A})\Delta) \\
 & + \alpha^2 \frac{\cos^2 \theta_{23} \sin^2 2\theta_{12}}{\hat{A}^2} \sin^2(\hat{A}\Delta)
 \end{aligned}$$

$$J_{CP} = 1/8 \sin \delta_{CP} \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23}$$

$$I_{CP} = 1/8 \cos \delta_{CP} \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23}$$

$$\alpha = \Delta m_{21}^2 / \Delta m_{31}^2, \quad \Delta = \Delta m_{31}^2 L / 4E$$

$$\hat{A} = 2VE / \Delta m_{31}^2 \approx (E_\nu / \text{GeV}) / 11 \quad \text{For earth's crust}$$

Electron neutrino appearance physics parameter extraction

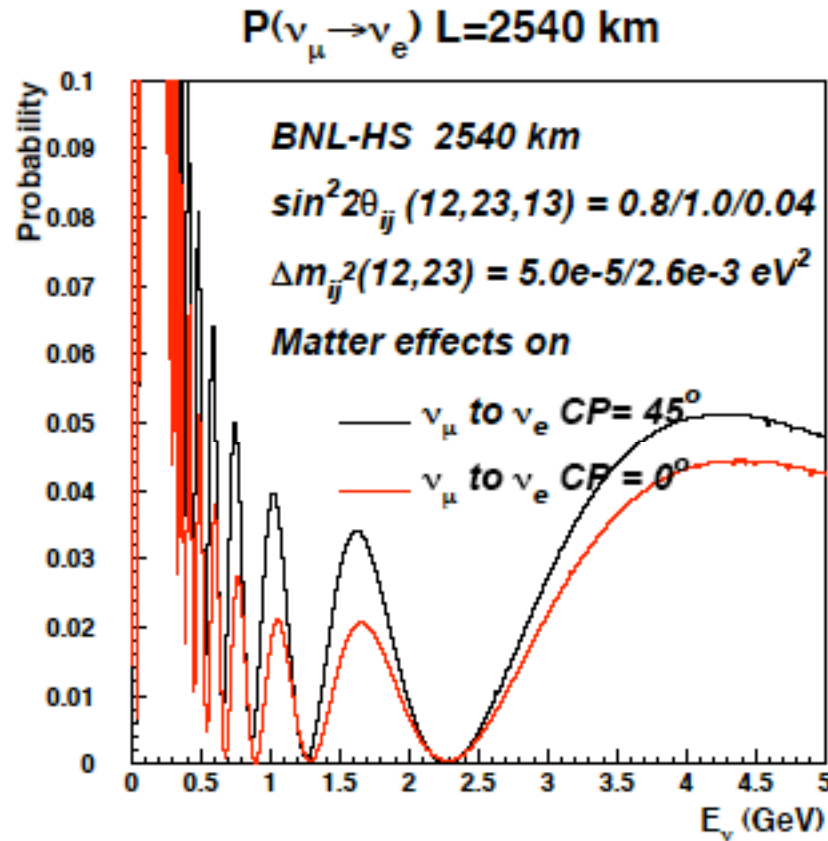
Assume $L > 2000\text{km}$, wide band beam

Δm_{32}^2 , Δm_{21}^2 , θ_{12} well known.

3 neutrino generations. $\uparrow\uparrow$ = large change \uparrow = small change

		$\sin^2 2\theta_{13} > 0$	$\Delta m_{32}^2 (> 0, < 0)$	$\delta_{CP} = (\pi/4, -\pi/4)$	$\theta_{23} (< \pi/4, > \pi/4)$
ν	0 – 1.2 GeV	\uparrow	–, –	\uparrow, \downarrow	$\uparrow\uparrow, \downarrow\downarrow$
	1.2 – 2.2 GeV	\uparrow	–, –	$\uparrow\uparrow, \downarrow\downarrow$	\downarrow, \uparrow
	> 2.2 GeV	\uparrow	$\uparrow\uparrow, \downarrow\downarrow$	\uparrow, \downarrow	\downarrow, \uparrow
$\bar{\nu}$	0 – 1.2 GeV	\uparrow	–, –	\downarrow, \uparrow	$\uparrow\uparrow, \downarrow\downarrow$
	1.2 – 2.2 GeV	\uparrow	–, –	$\downarrow\downarrow, \uparrow\uparrow$	\downarrow, \uparrow
	> 2.2 GeV	\uparrow	$\downarrow\downarrow, \uparrow\uparrow$	\downarrow, \uparrow	\downarrow, \uparrow

Numerical calculation



General Features

- 0.5 – 1 GeV: Δm_{12}^2 (LMA) region.
- 1 – 3 GeV: CP large effects region
- > 3 GeV: Matter enhanced (ν_μ), suppressed ($\bar{\nu}_\mu$). ($\Delta m_{32}^2 > 0$) Region.

Exact numerical calculation

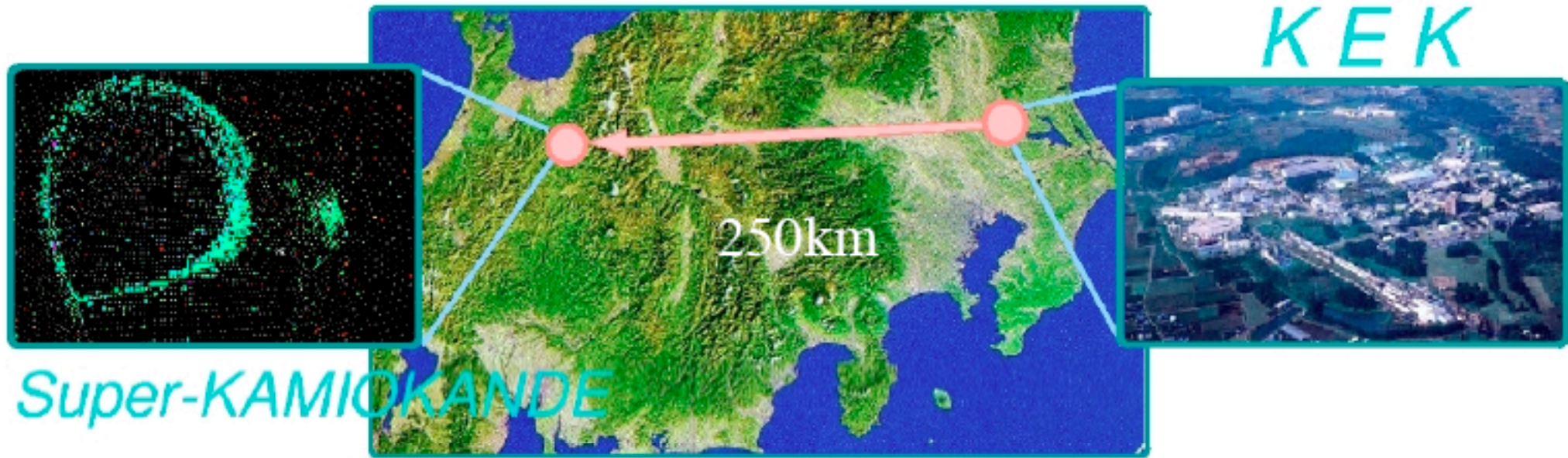
Simple rules

- Multiples nodes important for precision and new physics.
- Long distances separate CP and matter effects.
- Need $2500 \text{ kT} * \text{MW} * (10^7 \text{ sec})$ for measuring CP (regardless of distance and value of θ_{13})
- For CP violation study NO conventional beam experiment can get below $\sin^2 2\theta_{13} \sim 0.01$

New Age of Accelerator Neutrinos

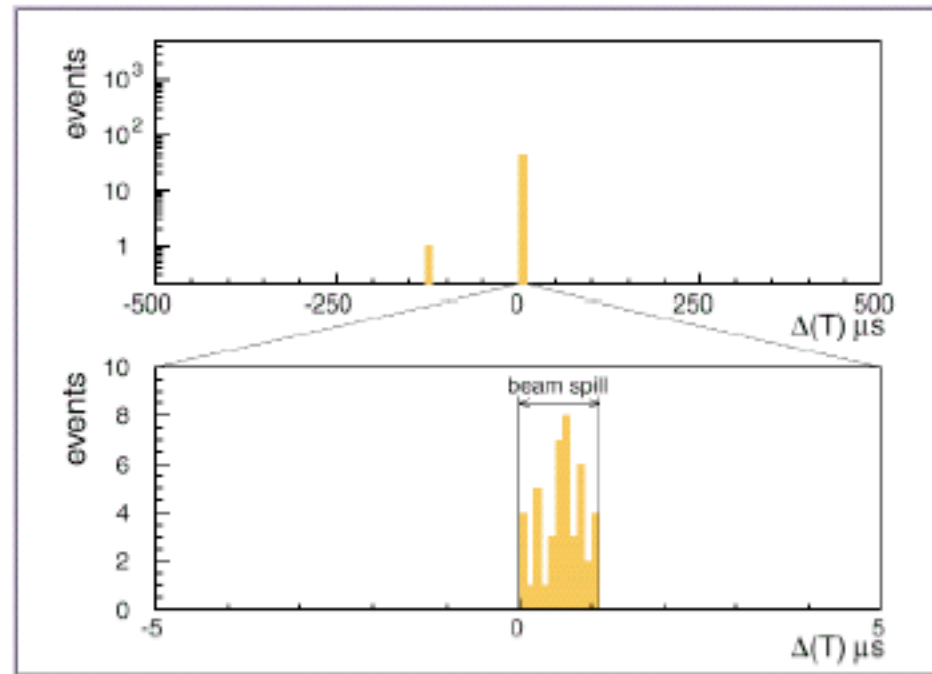
- For more precise experiments need pure beams of muon type neutrinos (or anti-neutrinos)
- Better controlled characteristics: energy, spectrum, backgrounds, pulsed.
- High energy (> 1 GeV) to provide events with long muons. Better resolution.
- Generally called Long Baseline Experiments.

Long Baseline Experiments

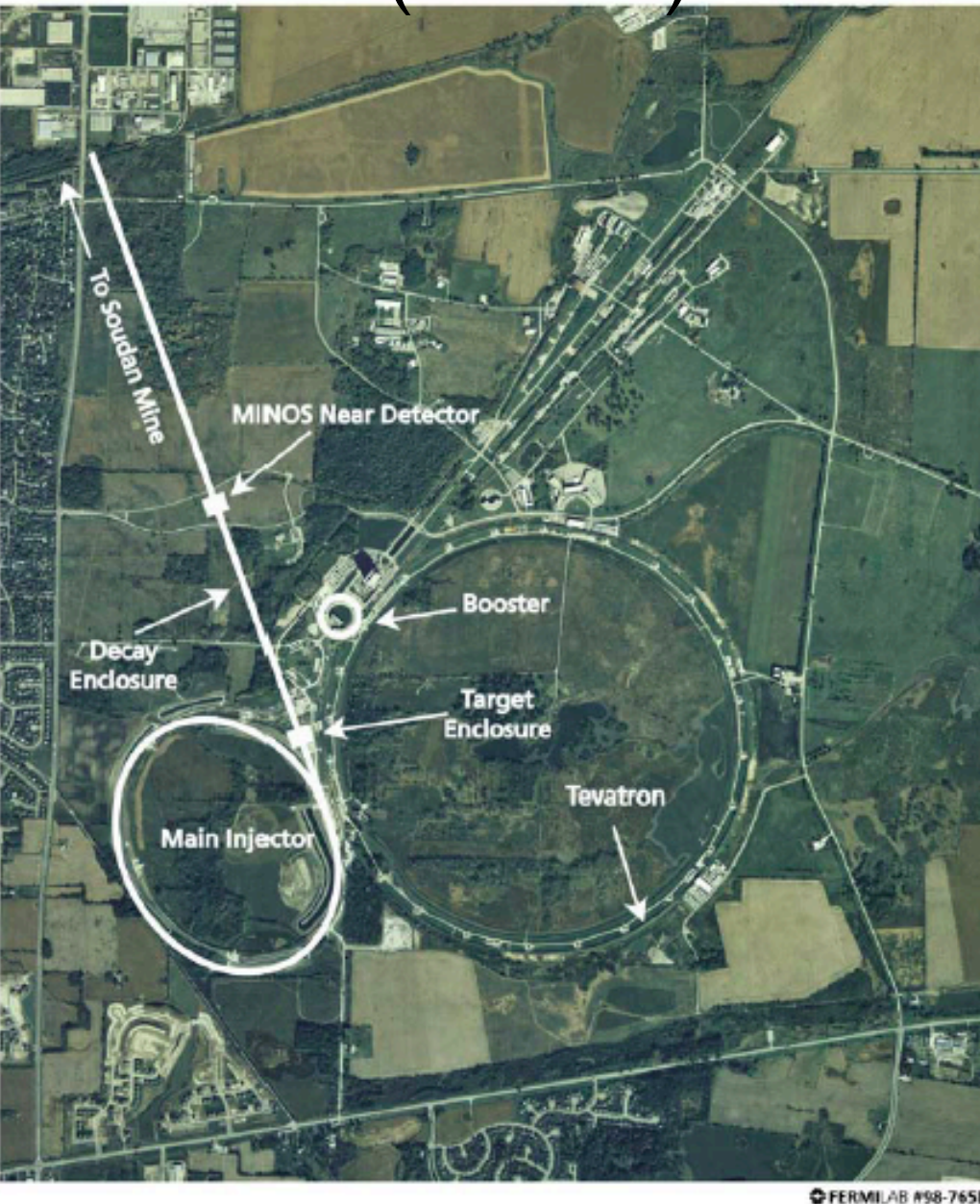


First LBL exp. with
positive result

81 ± 8 events no oscillation
56 events observed

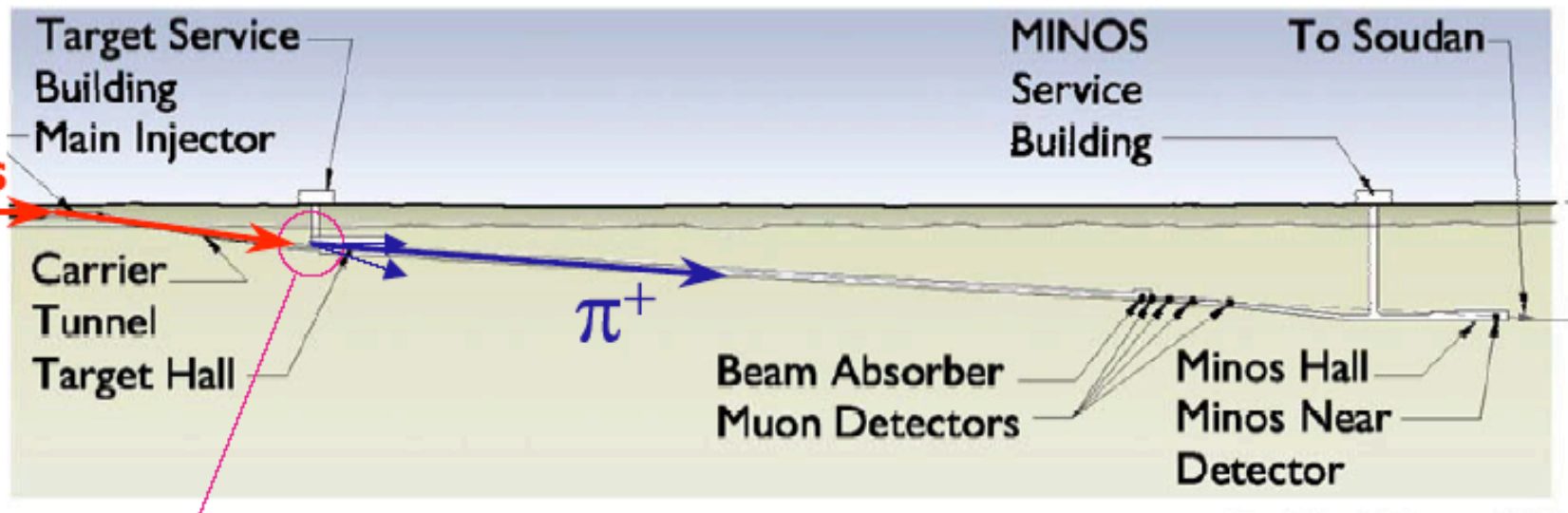


(Fermilab) Main Injector Neutrino Oscillation (MINOS) about to start running.



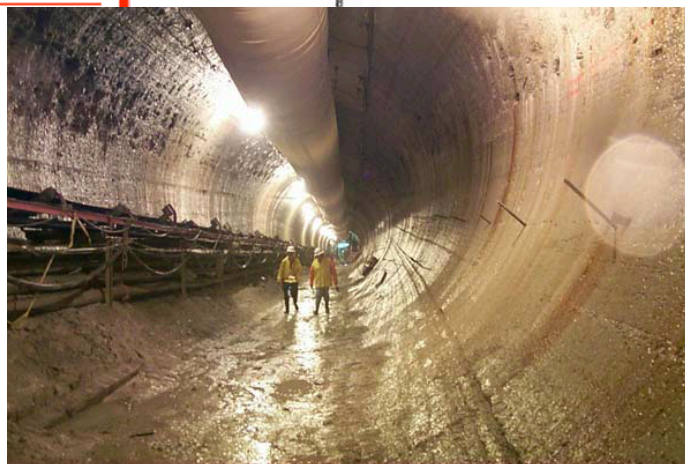
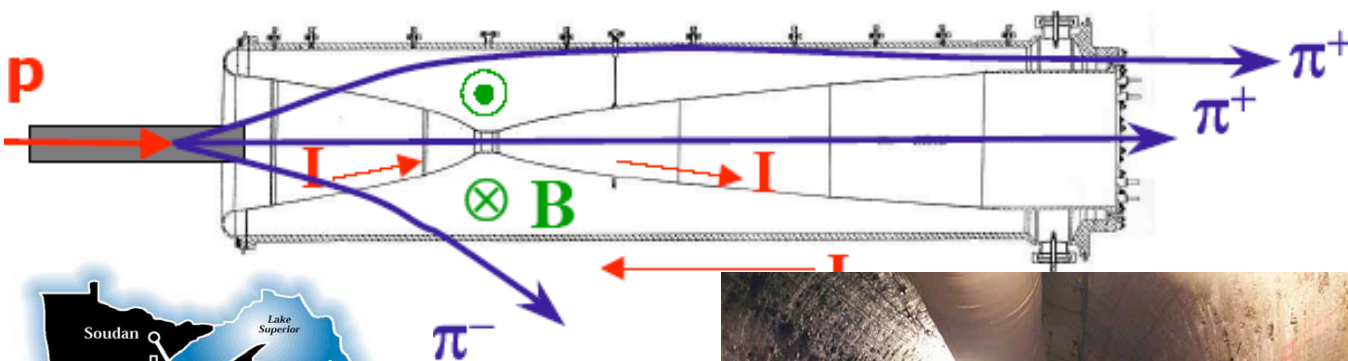
- ★ **120 GeV protons extracted from the MAIN INJECTOR in a single turn ($8.7\mu\text{s}$)**
- ★ **1.9 s cycle time**
- ★ ***i.e.* ν beam 'on' for $8.7\mu\text{s}$ every 1.9 s**
- ★ **2.5×10^{13} protons/pulse**
- ★ **0.3 MW on target !**
- ★ **Initial intensity**
 2.5×10^{20} protons/year

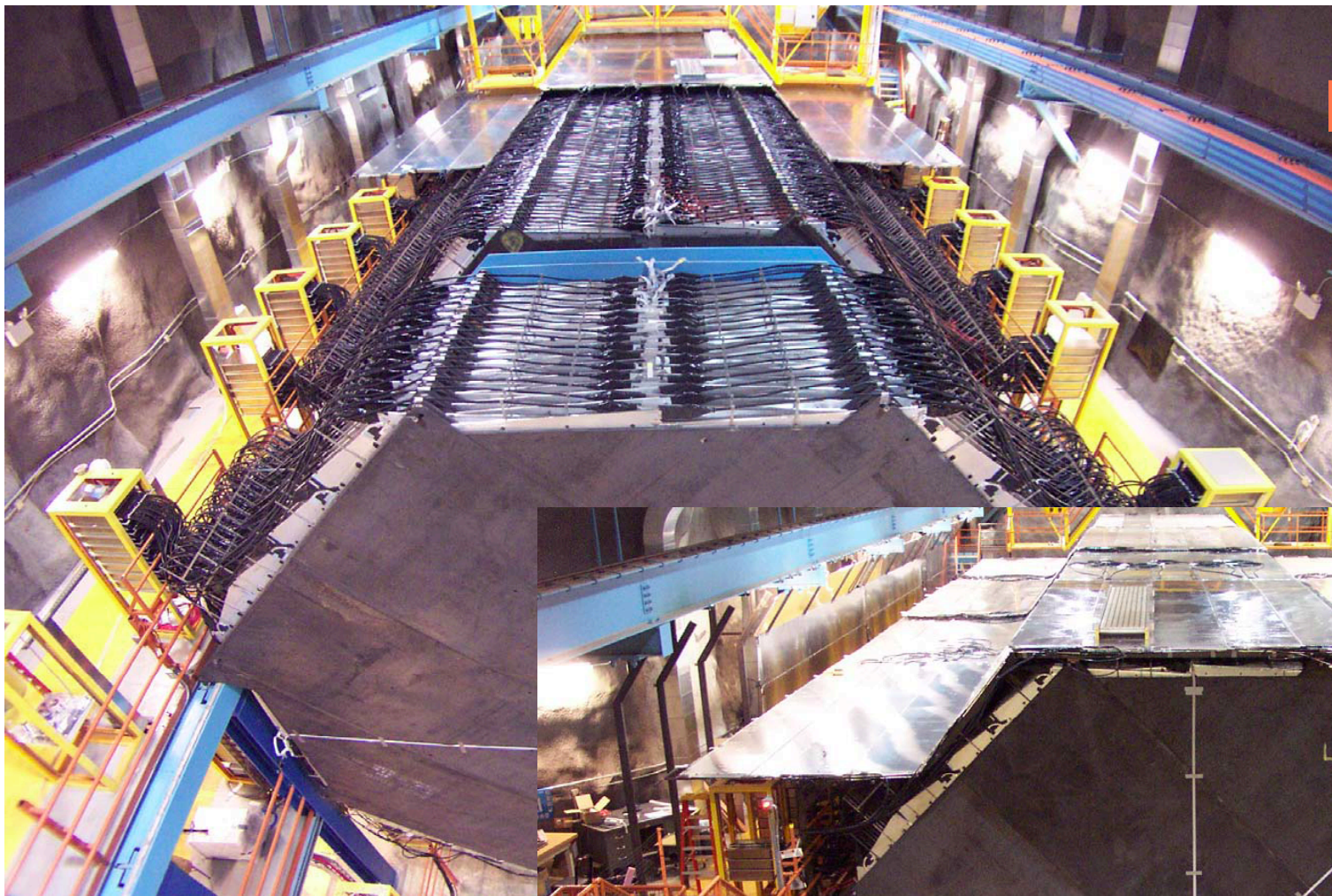
protons



0 64 128 256
METERS

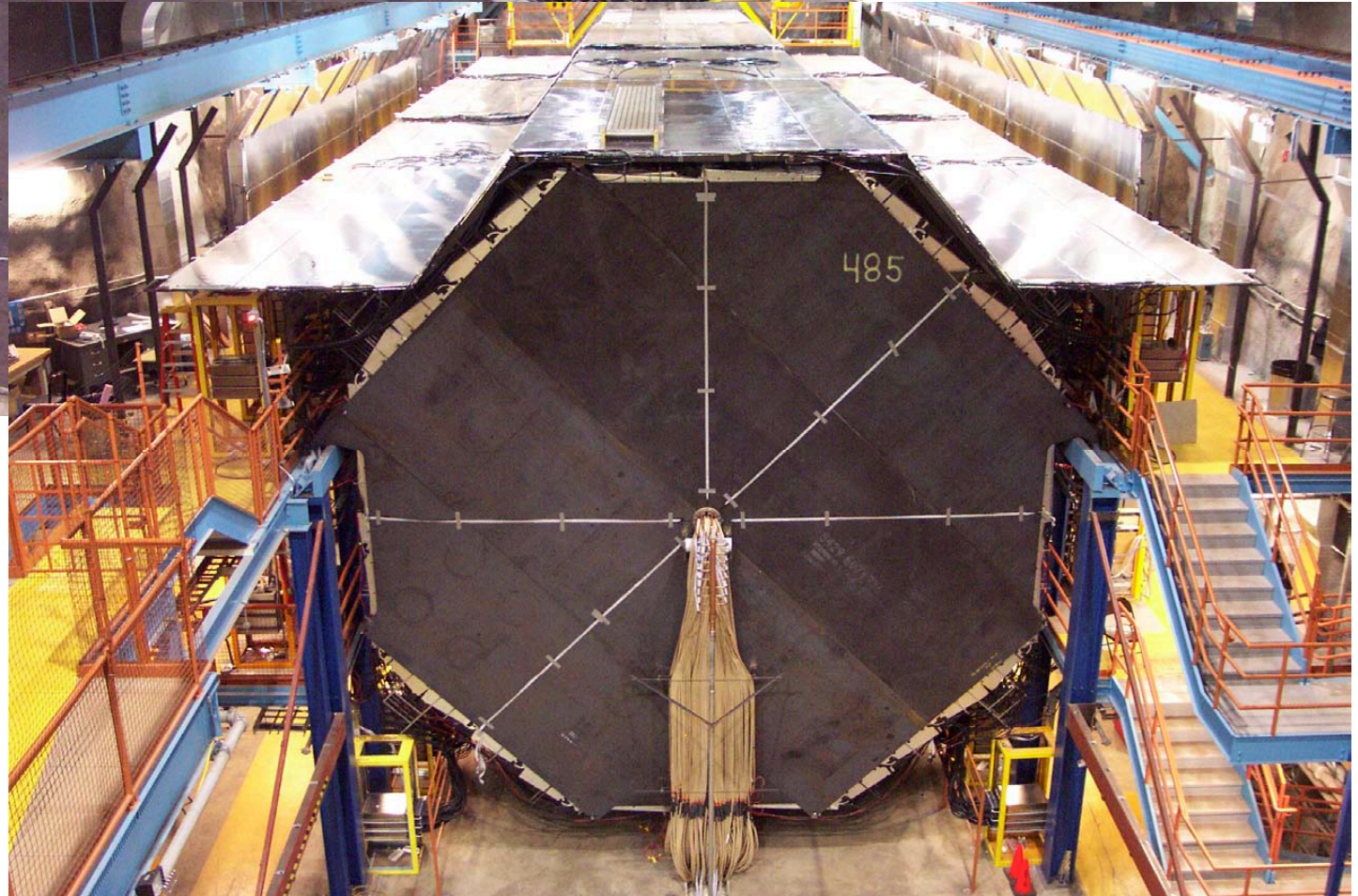
- Horn pulsed with 200 kA
- Toroidal Magnetic field $B \sim I/r$ between inner and outer conductors



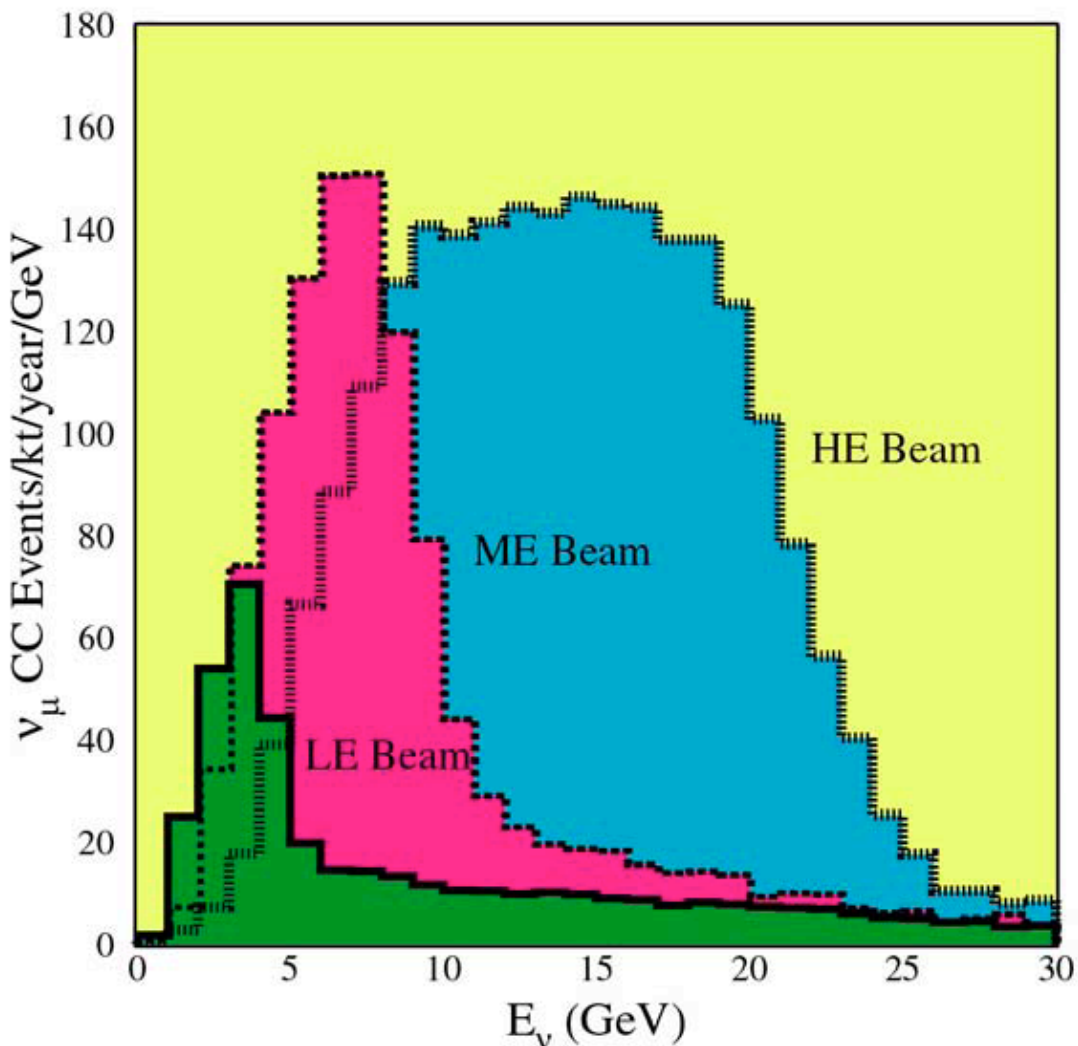


Fully operational
in Soudan mine
at 2341 ft
730 km from
FNAL

Minos
detector:
Iron/
scintillator
5kT



MINOS Physics Plots



LE BEAM:

ν_μ CC Events/year:

Low

Medium

High

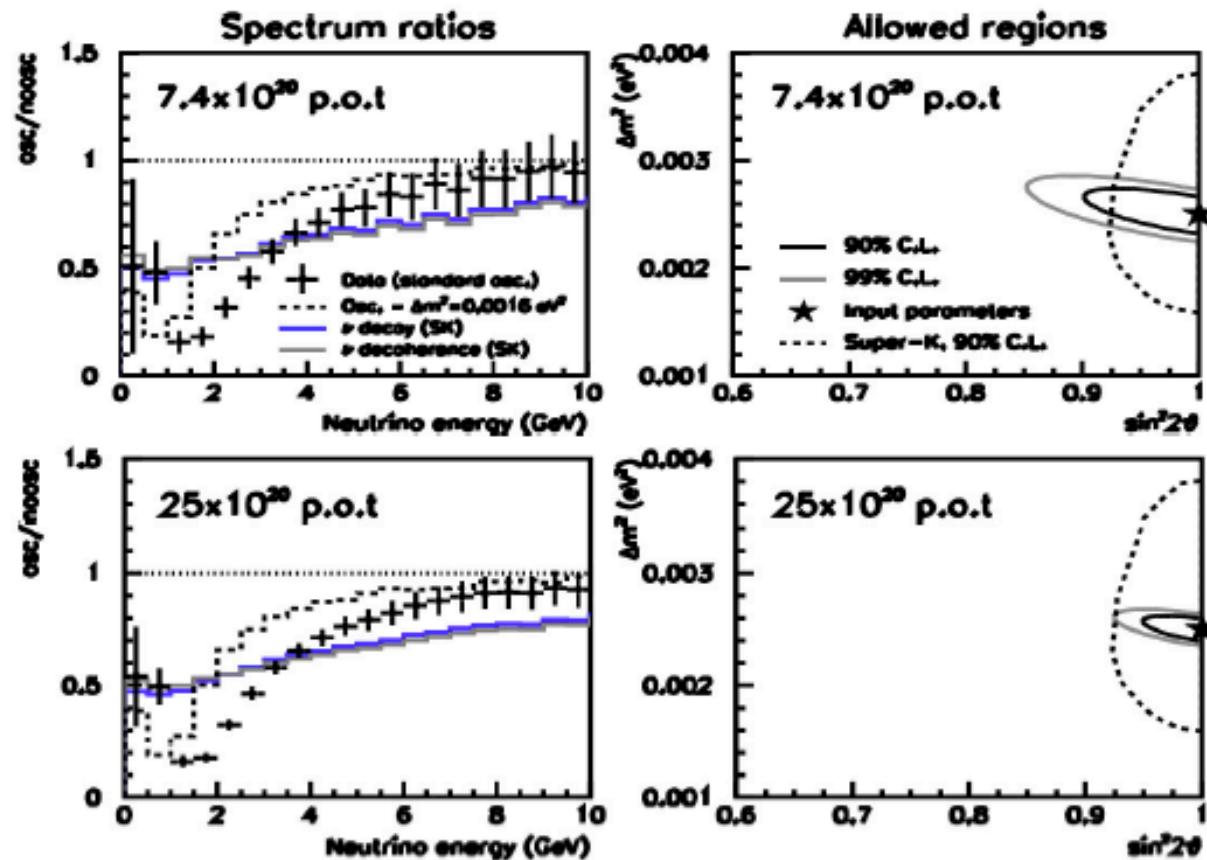
1600

4300

9250

(2.5×10^{20} protons on target/year)

★ Measurement of Δm^2 and $\sin^2 2\theta$



For $\Delta m^2 = 0.0025 \text{ eV}^2$,
 $\sin^2 2\theta = 1.0$

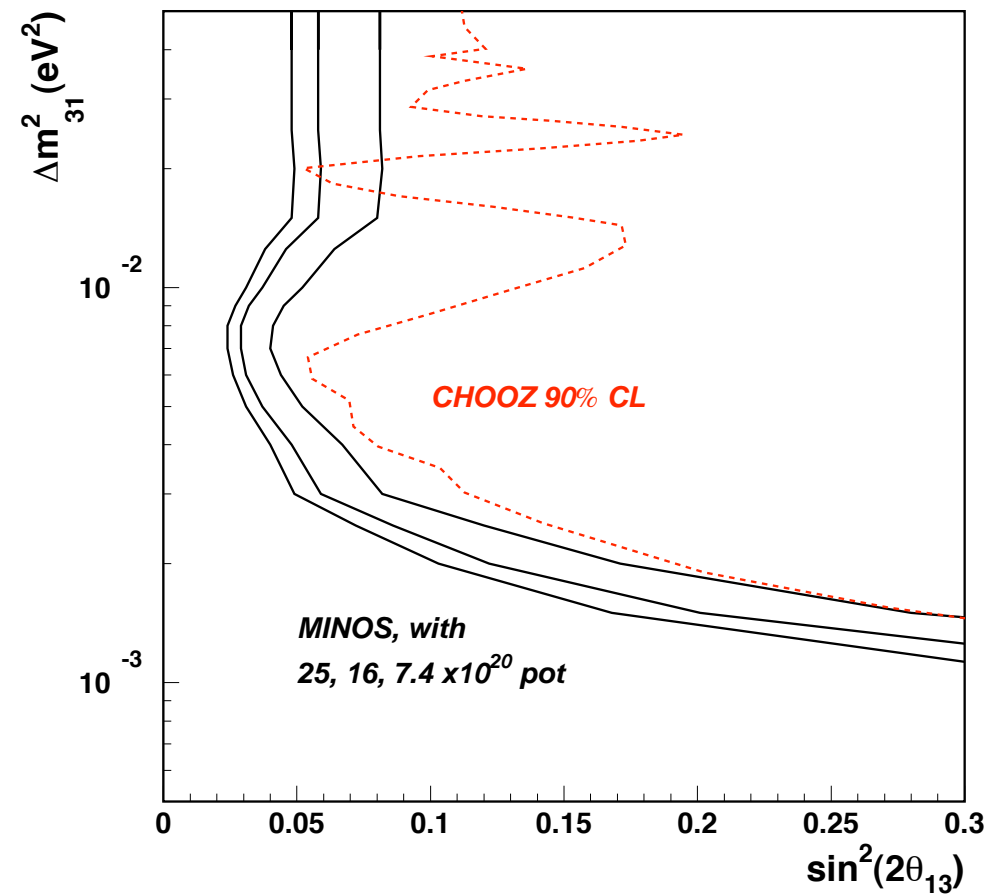
Large improvement in
precision !

Final sensitivity depends
on protons on target

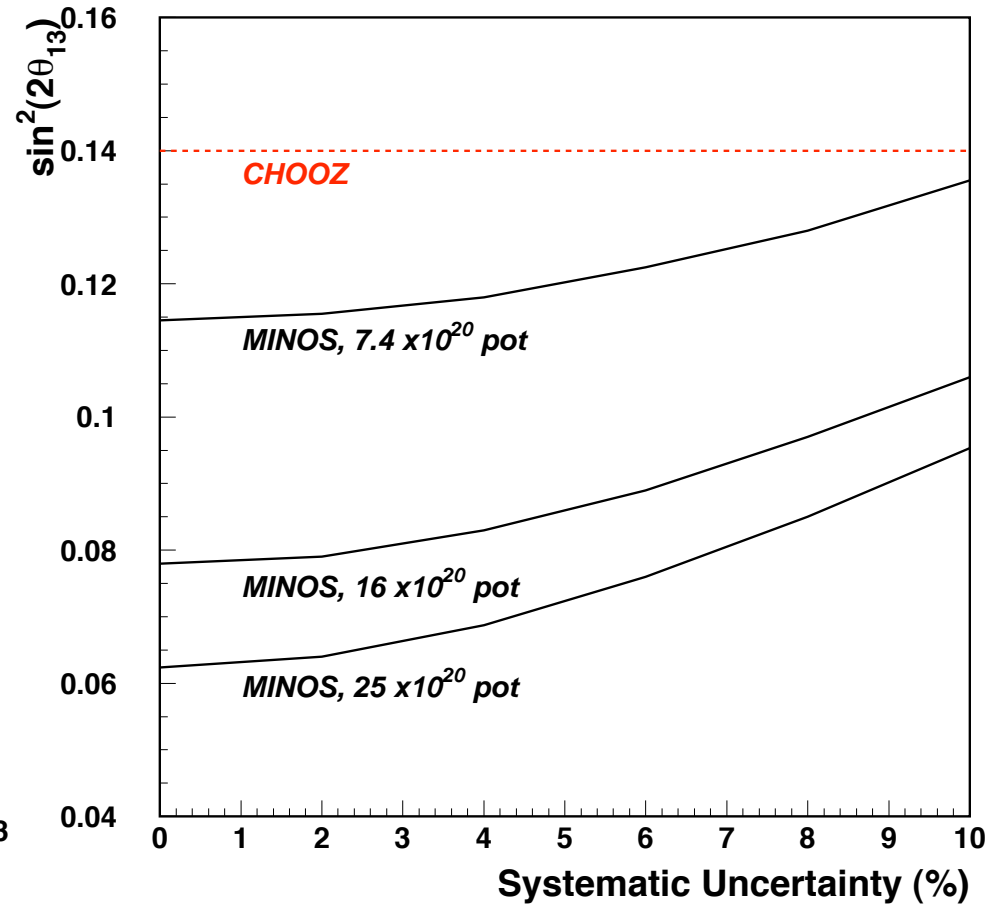
- ★ Direct measurement of **L/E** dependence of ν_μ flux
- ★ Powerful test of flavour oscillations vs. alternative models

Measurement of θ_{13} in Minos

3 σ Contours

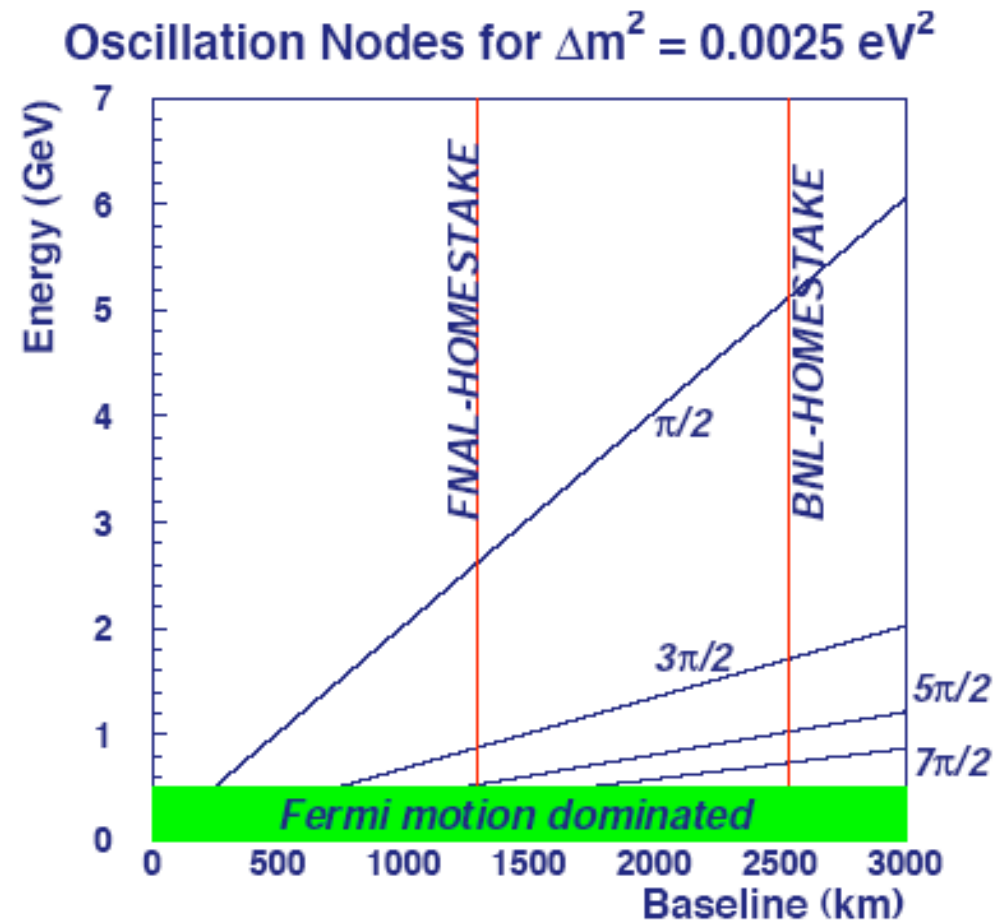


3 σ Discovery Potential



Ultimate Ambitions !

- Must see multiple nodes in a spectrum for precise measurements
- Need E: 1-6 GeV
- Need ~ 2000 km
- Need intense beam.
- Need very large detector.
- PRD 68 (2003)012002

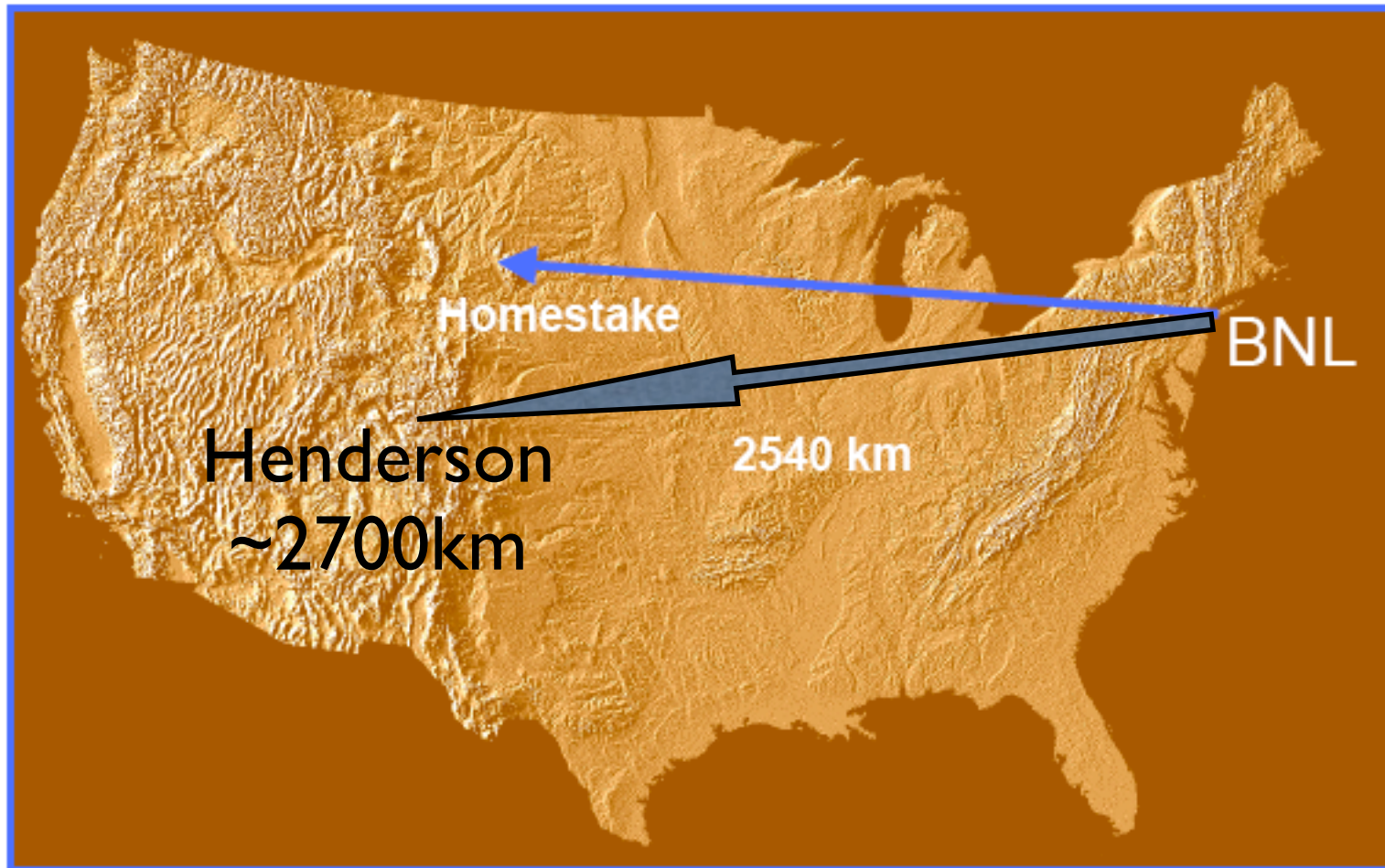


(M. Diwan, hep-ex/0407047)

My biases

- New detector facility (100-500kT) must address broad range of physics: Accelerator Neutrinos, Nucleon decay, Astrophysical Neutrinos.
- New detector most likely located deep.
- For neutrino oscillations must address the most difficult problem for reasonable range of parameters: CP violation.
- Exciting opportunity for younger people to make their mark.
- Does this make sense from physics point of view ?

BNL → Homestake 1 MW Neutrino Beam



28 GeV protons, 1 MW beam power

500 kT Water Cherenkov detector

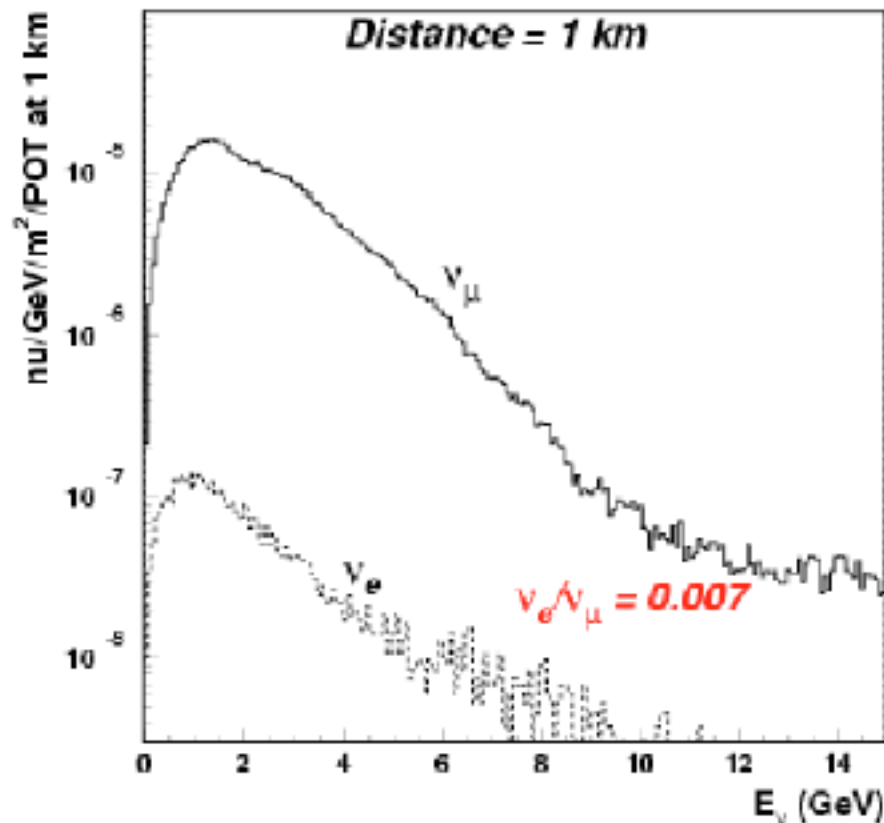
5e7 sec of running, Conventional Horn based beam

Update on AGS based Super Neutrino beam

- New conceptual design document
BNL-73210-2004-IR. (sent to DOE)
- http://raparia.sns.bnl.gov/nwd_ad
- Redesigned beam facility: more compact,
now possible to make decay pipe longer.
- Completely new design for injector LINAC:
cheaper and faster to build.

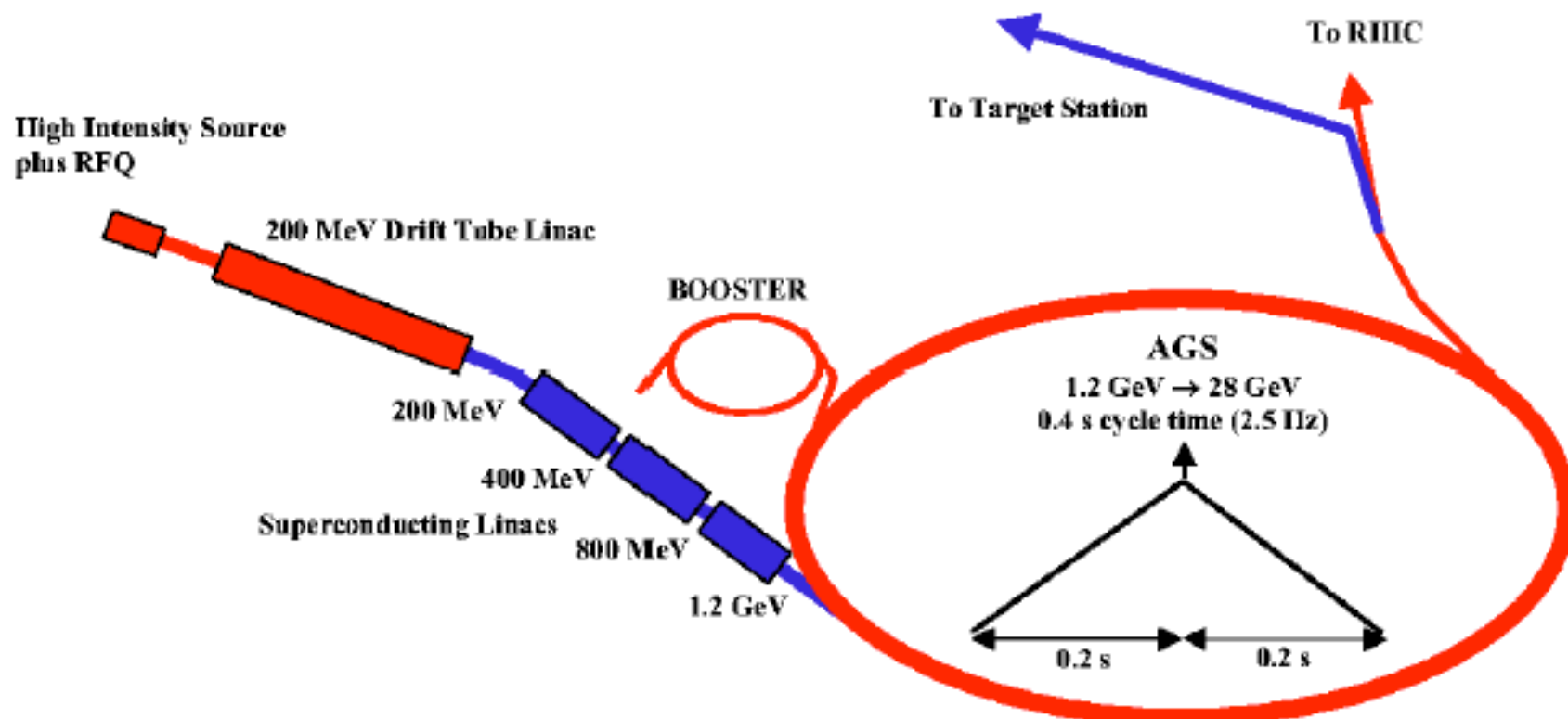
Neutrino spectrum from AGS

BNL Wide Band. Proton Energy = 28 GeV



- Proton energy 28 GeV
- 1 MW total power
- $\sim 10^{14}$ proton per pulse
- Cycle 2.5 Hz
- Pulse width 2.5 μ s
- Horn focused beam with graphite target
- 5×10^{-5} $\nu/\text{m}^2/\text{POT}$ @ 1km
- 52000 CC events.
- 17000 NC events.

BNL-AGS Target Power Upgrade to 1 MW

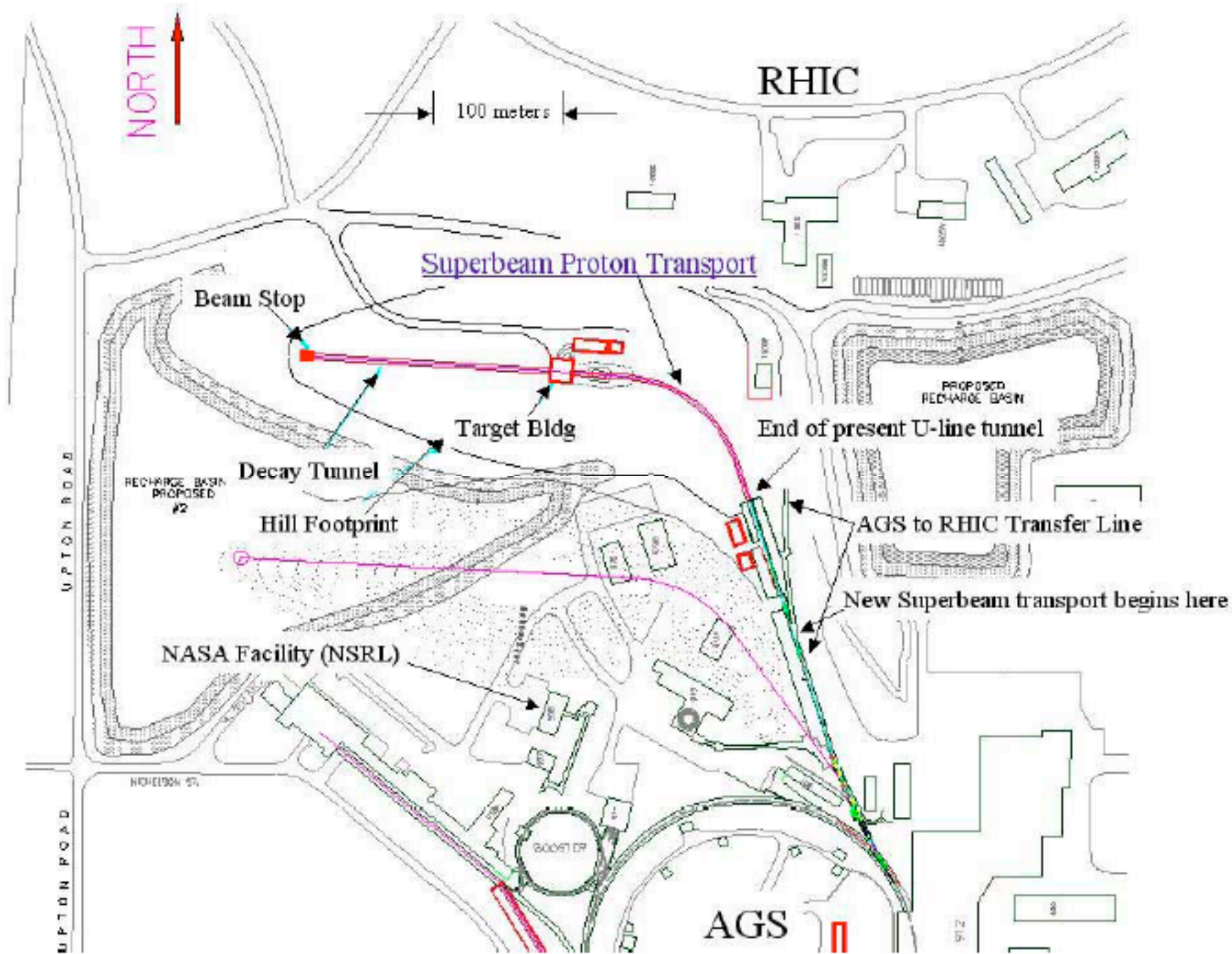


AGS is currently the highest intensity machine.
Simple plan. Run the AGS faster. 2.5 Hz
Need new LINAC @ 1.2 GeV to provide
protons.

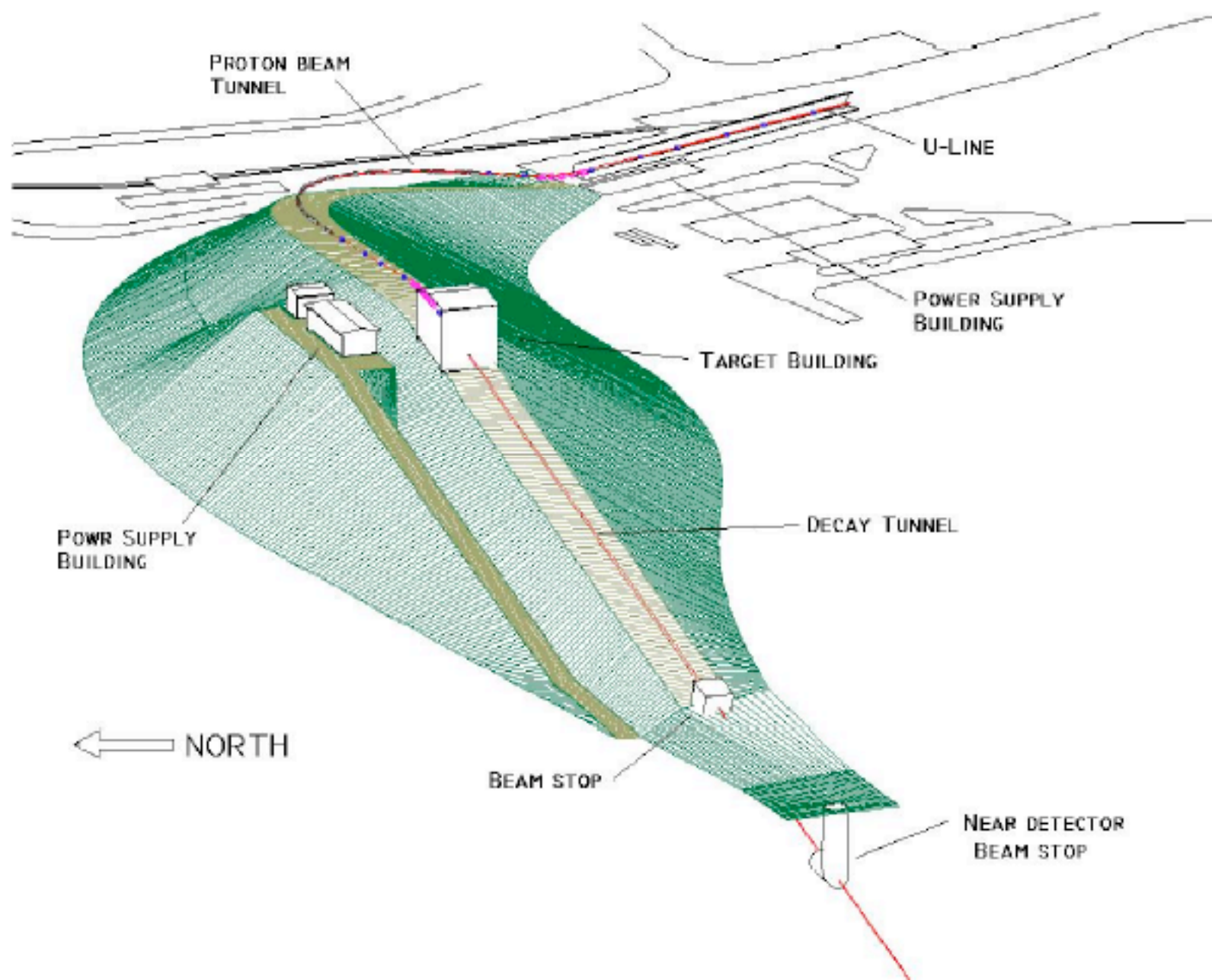
Cost \$265M FY03 (TEC) dollars.

Energy is 28 GeV. 2.5 Hz operation is 1 MW

$$7 \times 10^{13} \text{ protons}/2\text{sec}$$
$$9 \times 10^{13} \text{ protons}/0.4\text{sec}$$



3-D Neutrino Super Beam Perspective



Cost Estimate of the AGS Super Neutrino Beam Facility

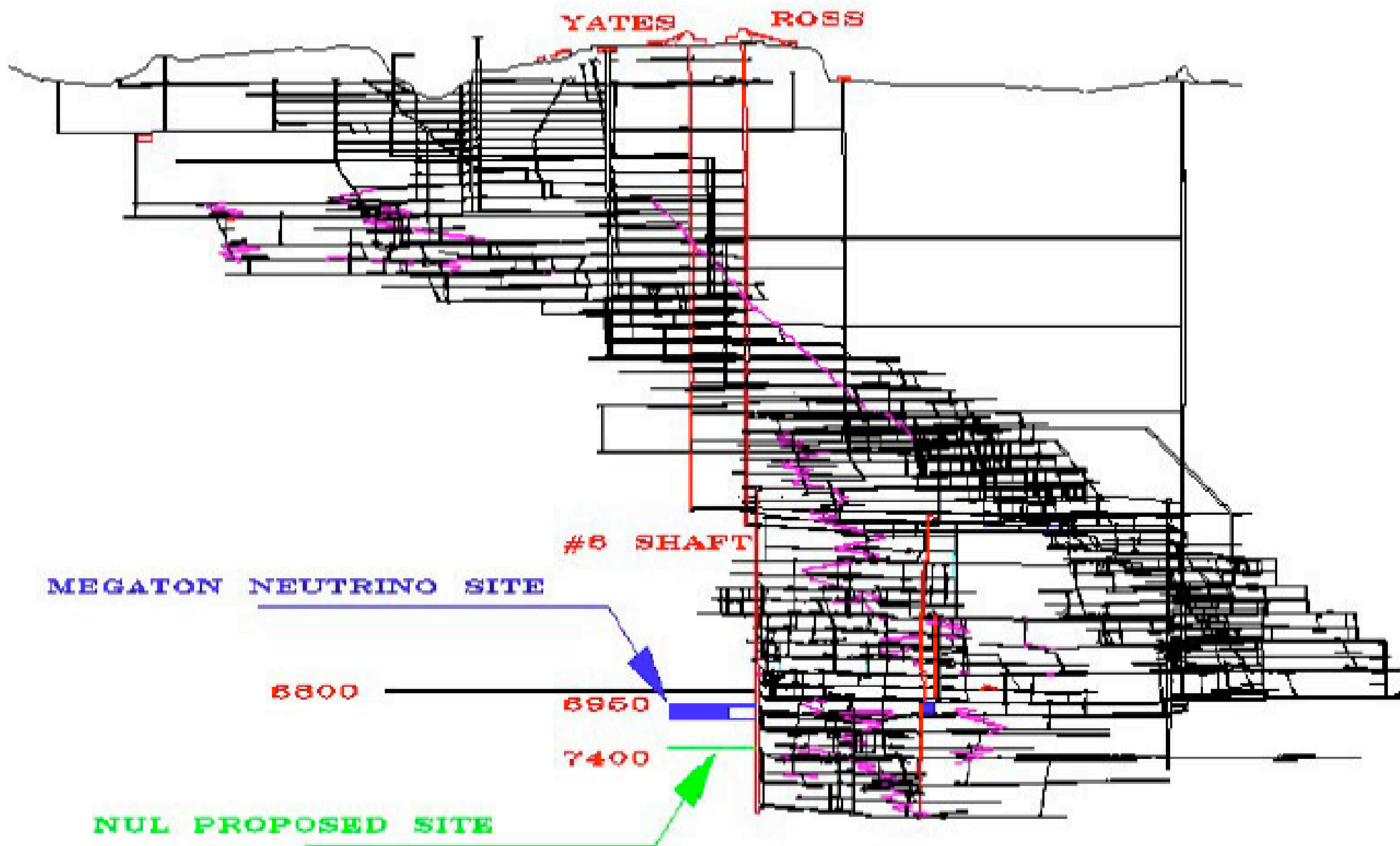
Construction Phase - Direct FY04 Dollars

1.0 AGS Super Neutrino Beam Facility	EDIA	M&S	Labor	Total
1.1 The Linac System	6,879,116	98,556,970	16,783,762	122,219,848
1.1.1 Front End and RT Linac Upgrade	313,000	2,383,000	856,000	3,552,000
1.1.2 SCL Accelerating Cavity System	954,240	22,254,200	11,040,000	34,248,440
1.1.3 SCL RF Source	3,620,998	51,668,800	402,332	55,692,120
1.1.4 SCL Cryogenic System	370,000	13,700,000	2,200,000	16,270,000
1.1.5 SCL Vacuum System	641,598	3,474,570	1,148,378	5,264,546
1.1.6 SCL Instrumentation	460,957	1,390,400	409,061	2,260,418
1.1.7 SCL Magnet and Power Supply	518,332	3,686,000	727,991	4,932,324
1.2 The AGS Upgrade	10,496,245	53,619,159	6,472,590	70,587,994
1.2.1 AGS Main Magnet Power Supply	503,959	28,200,000	1,342,337	30,046,296
1.2.2 AGS RF System Upgrade	6,082,625	9,850,000	675,847	16,608,472
1.2.3 AGS Injection/Extraction	644,000	6,437,068	1,668,330	8,749,396
1.2.4 Beam Transport to Target	1,636,771	7,852,241	2,637,290	12,126,302
1.2.5 Control System	1,628,890	1,279,852	148,786	3,057,528
1.3 The Target and Horn System	664,742	3,417,152	1,208,338	5,290,232
1.3.1 The Target System	127,008	229,284	50,130	406,422
1.3.2 The Horn System	454,524	2,358,568	656,224	3,469,316
1.3.3 Shielding and Remote Handling	83,210	809,300	125,300	1,017,810
1.3.4 Target & Horn Physics Support	0	20,000	376,684	396,684
1.4 The Conventional Facility	7,550,300	60,090,300	1,210,700	68,851,300
1.4.1 Linac Tunnel/Klystron Gallery	2,253,000	11,529,000	230,000	14,012,000
1.4.2 AGS Power Supply Building	2,024,000	13,347,000	432,000	15,803,000
1.4.3 Beam Transport and Target Area	1,674,300	25,091,000	172,500	26,937,800
1.4.4 The Decay Tunnel and Beam Stop	184,000	1,225,300	115,200	1,524,500
1.4.5 Site Utilities & Roads	1,088,000	6,820,000	140,000	8,048,000
1.4.6 Modifications for AGS RF System	327,000	2,078,000	121,000	2,526,000
1.5 ES&H	104,652	275,211	437,355	817,218
1.5.1 ES&H	20,000	105,000	270,000	395,000
1.5.2 Access Controls	84,652	170,211	167,355	422,218
1.6 Project Support	1,148,681	384,109	4,096,963	5,629,753
1.6.1 Project Management	0	100,000	1,178,000	1,278,000
1.6.2 Technical Support	1,148,681	214,109	2,146,963	3,509,753
1.6.3 Project Controls	0	70,000	772,000	842,000
AGS Super Neutrino Beam Facility Project Total	25,843,736	216,342,901	30,209,709	273,396,345

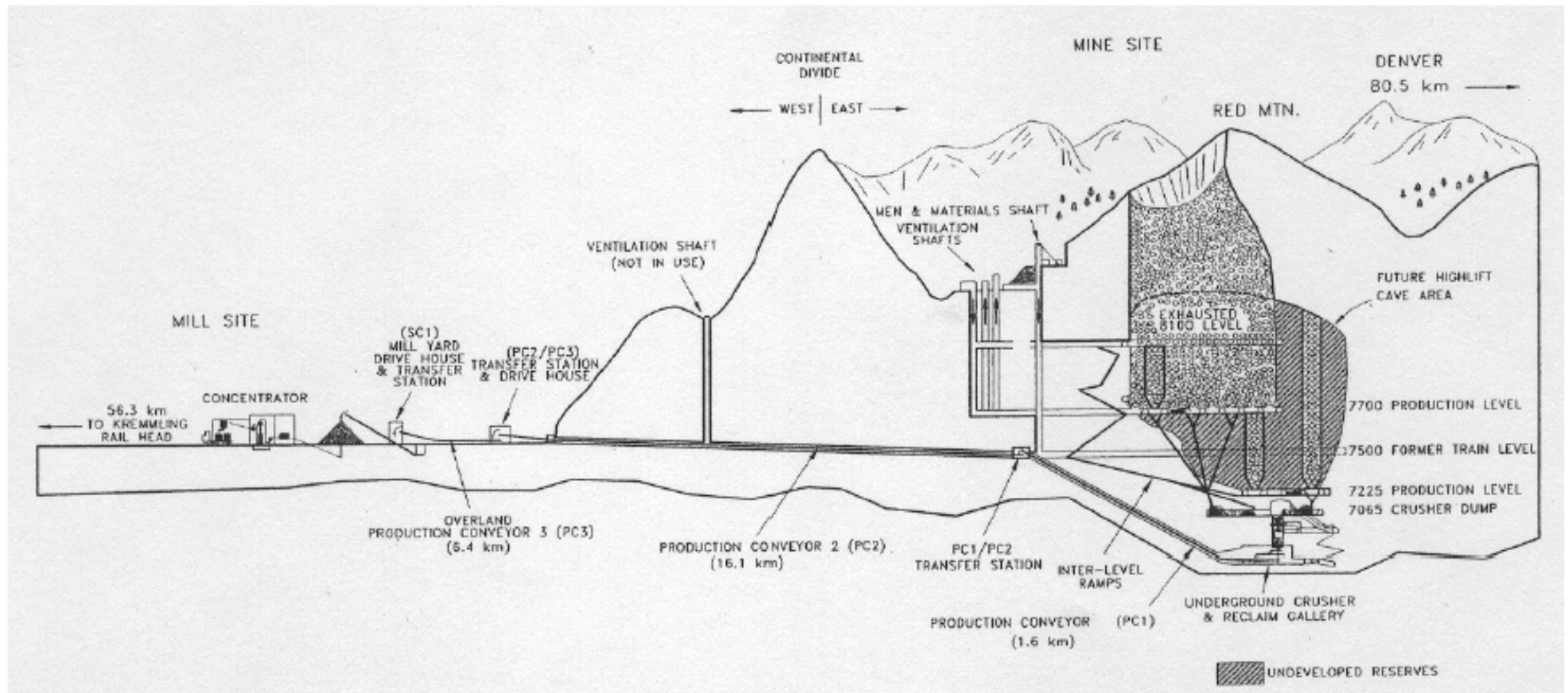
Deep Underground Laboratory Initiative

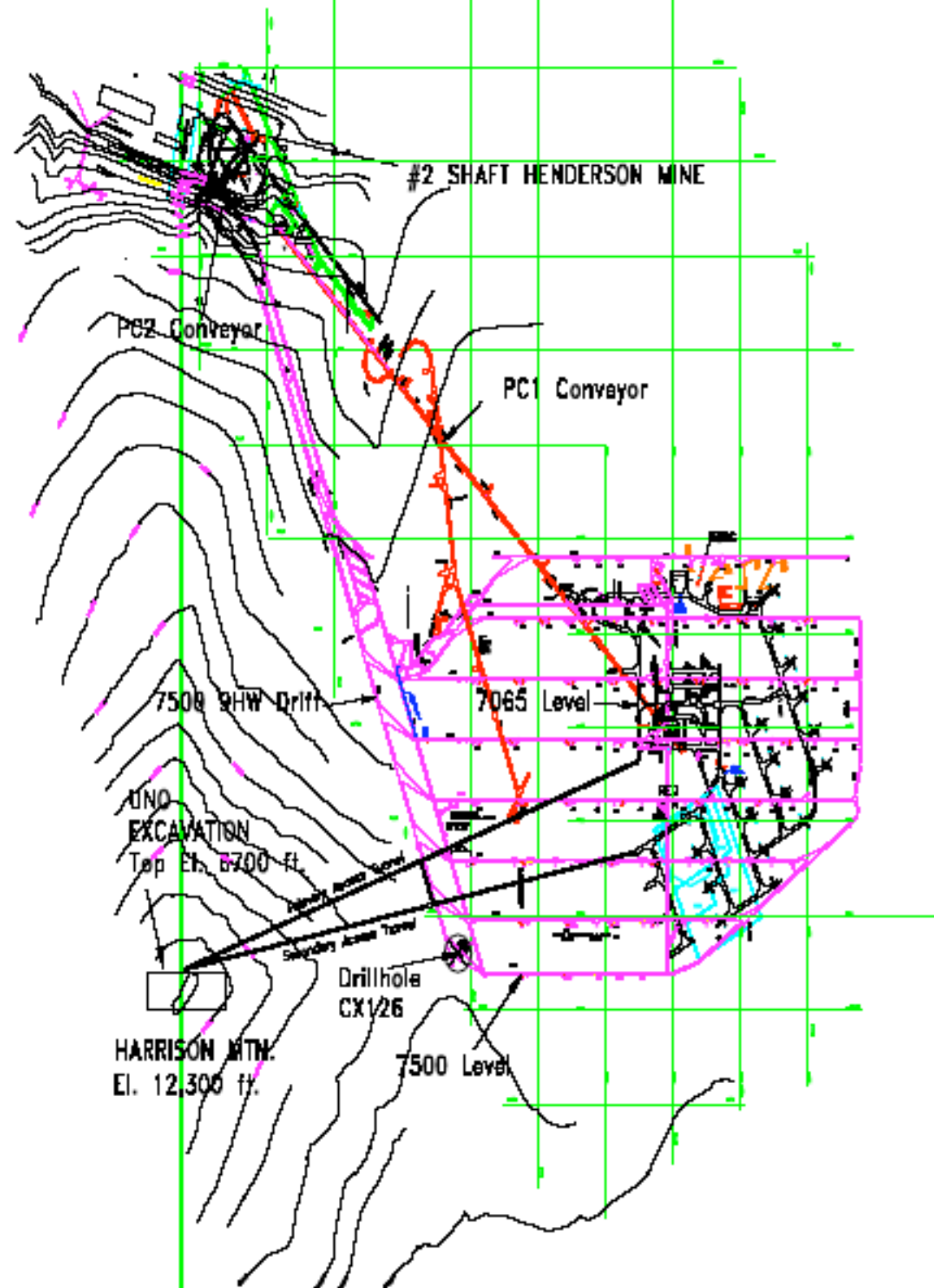
- New discussion started when Homestake gold mine (site of Davis Chlorine experiment) closed.
- National Science Foundation has initiated a series of solicitations.
- S1 - focusses on science first. Identify all science (physics, geology, biology) and infrastructure needs.
<http://neutrino.lbl.gov/DUSELS-I>
- S2 - decide on a suitable site.





Henderson mine





Detector

- Requirements: Very ambitious !
 - 500 kTons fiducial mass for both Proton decay and neutrino astro-physics and neutrino beam physics.
 - $\sim 10\%$ energy resolution on quasielastic events
 - Muon/electron discrimination at $< 1\%$
 - 1, 2, 3 track event separation
 - Showering NC event rejection at factor of ~ 15
 - Low threshold (~ 10 - 15 MeV) for supernova search
 - Part of the detector could have lower threshold for solar neutrino detection.
 - Time resolution of \sim few ns for pattern recognition and background reduction.

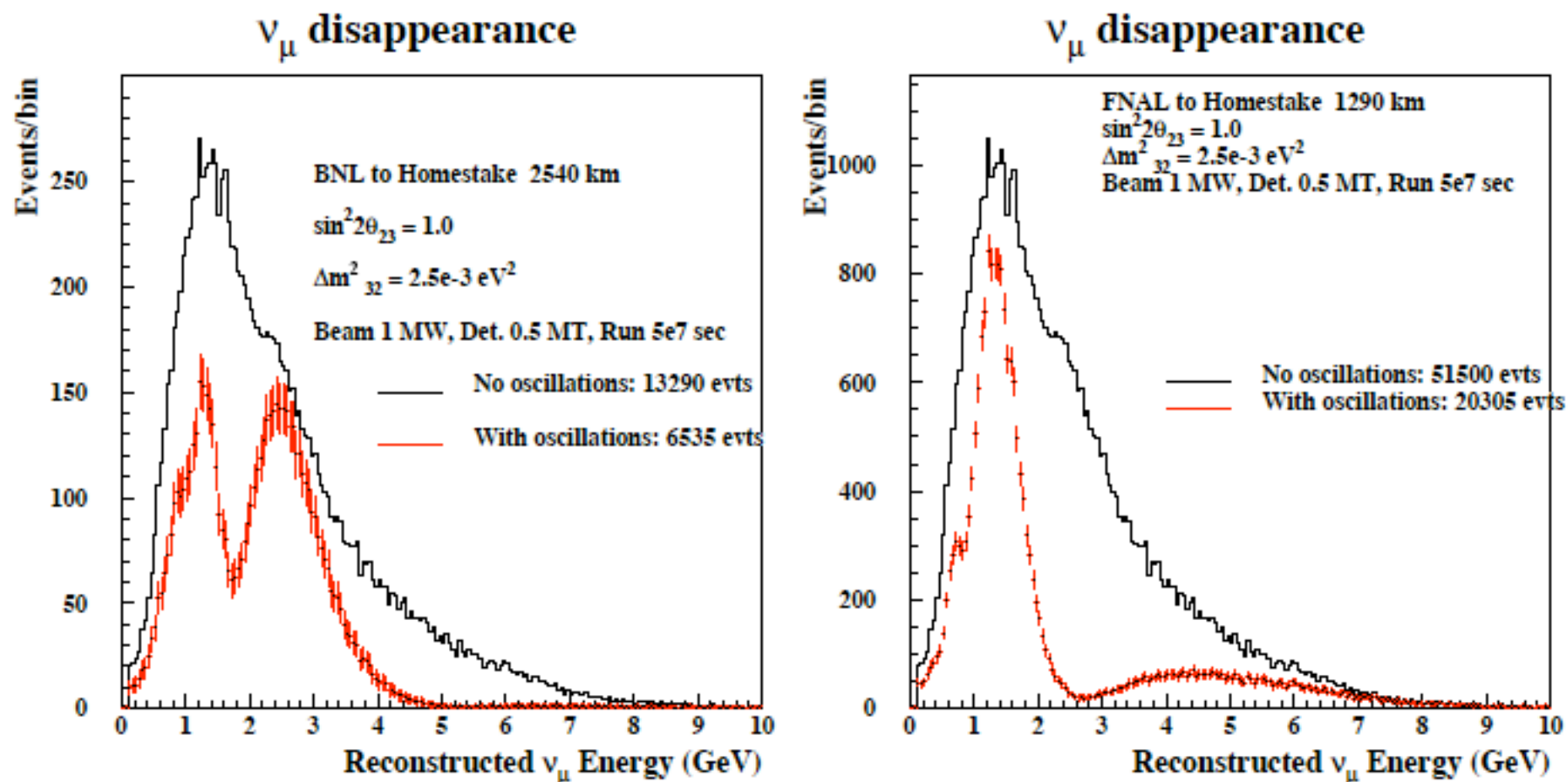


Figure 2: *Simulated spectrum of detected muon neutrinos for 1 MW beam and 500 kT detector exposed for 5×10^7 sec. Left side is for baseline of 2540 km, right side for baseline of 1290 km. The oscillation parameters assumed are shown in the figure. Only clean single muon events are assumed to be used for this measurement (see text).*

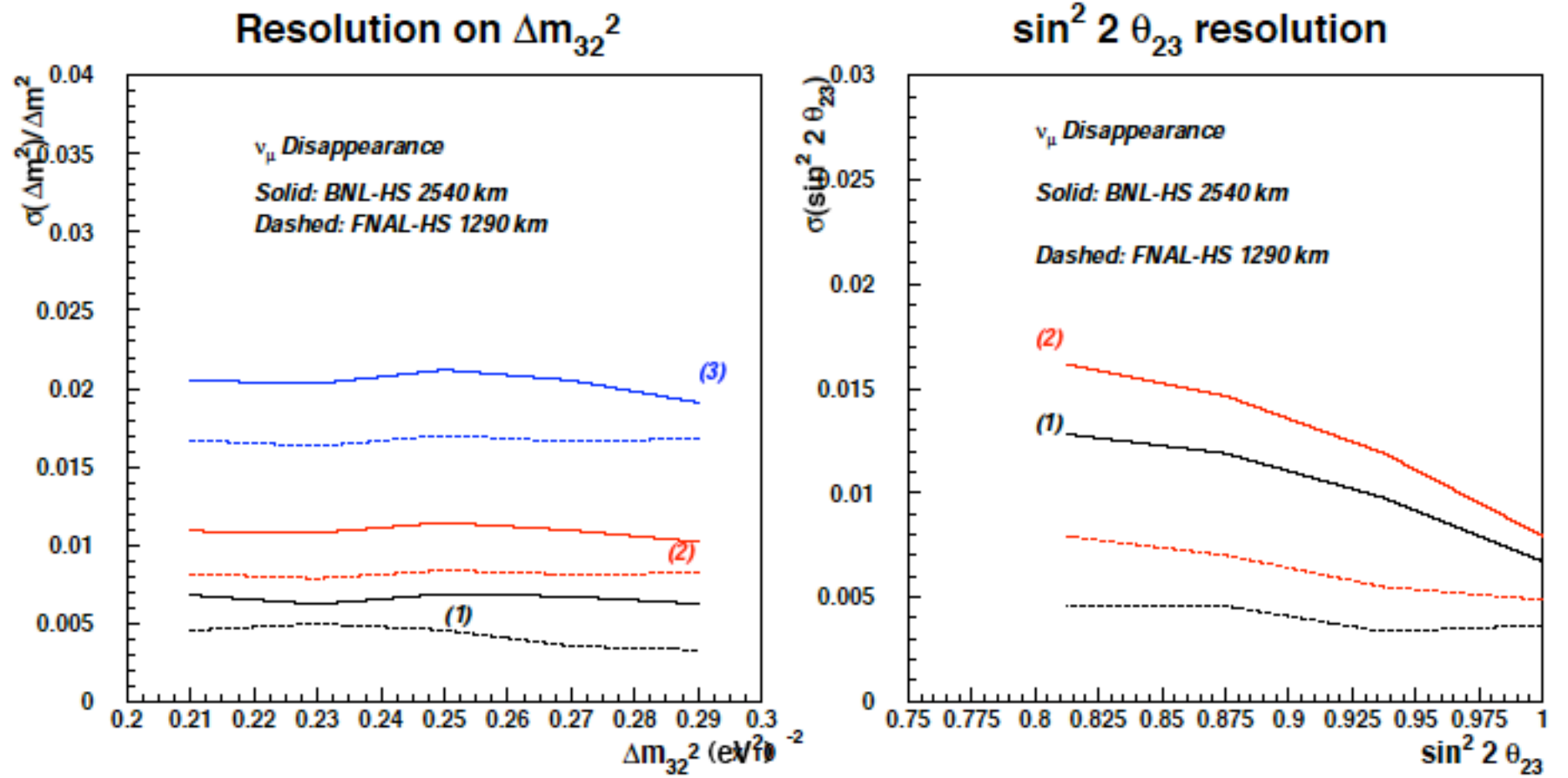
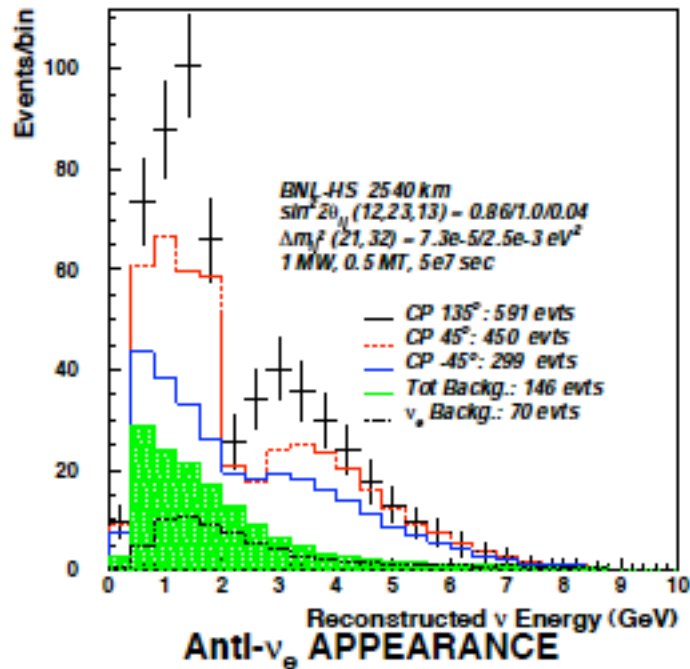
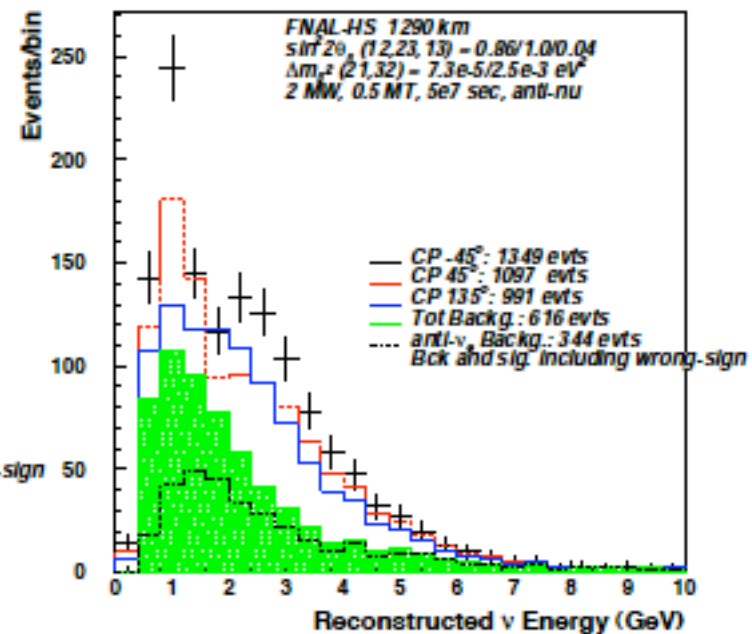
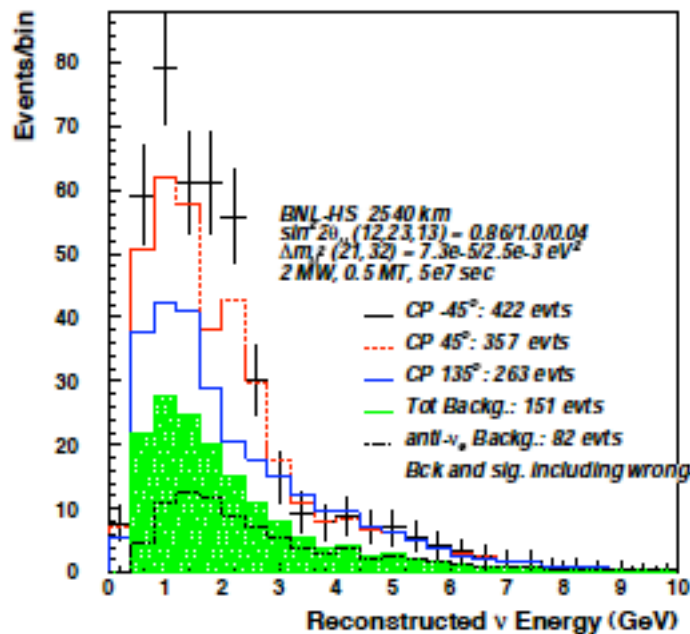
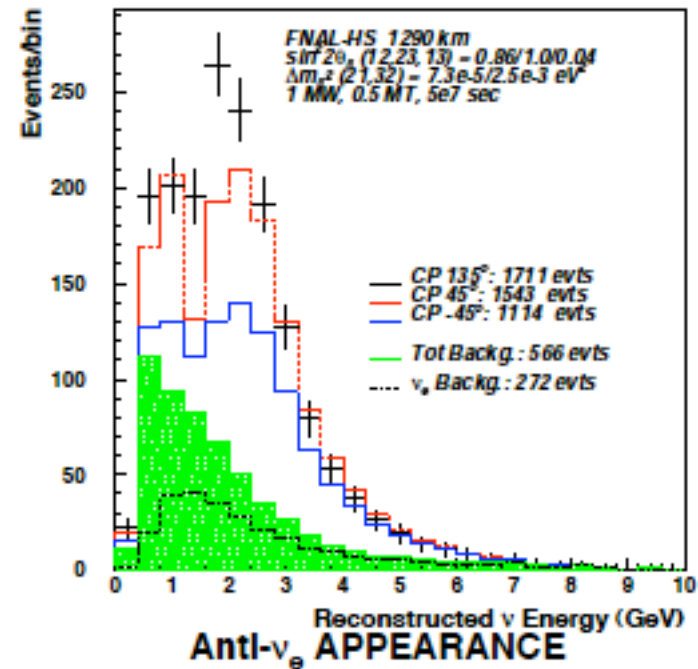


Figure 3: 1 sigma resolutions on Δm_{32}^2 (left) and $\sin^2 2\theta_{23}$ (right) expected after analysis of the oscillation spectra from Figure 2. The solid curves are for BNL-HS 2540 km baseline, and the dashed are for FNAL-HS 1290 km baseline. The curves labeled 1 and 2 correspond to statistics only and statistics and systematics, respectively (similarly for dashed curves of the same color). The curve labeled (3) on the left has an additional contribution of 1% systematic error on the global energy scale.

ν_e APPEARANCE



ν_e APPEARANCE



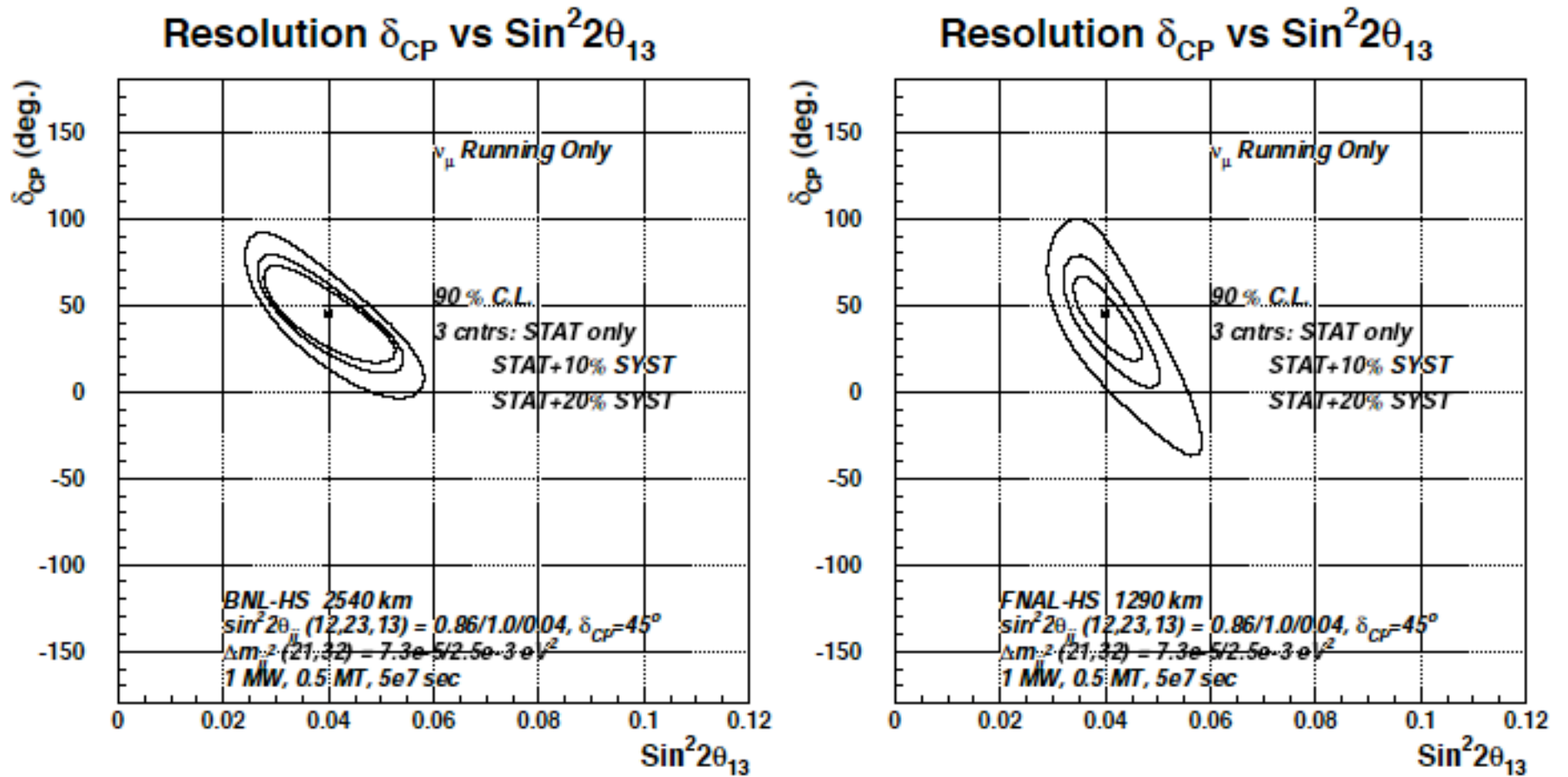
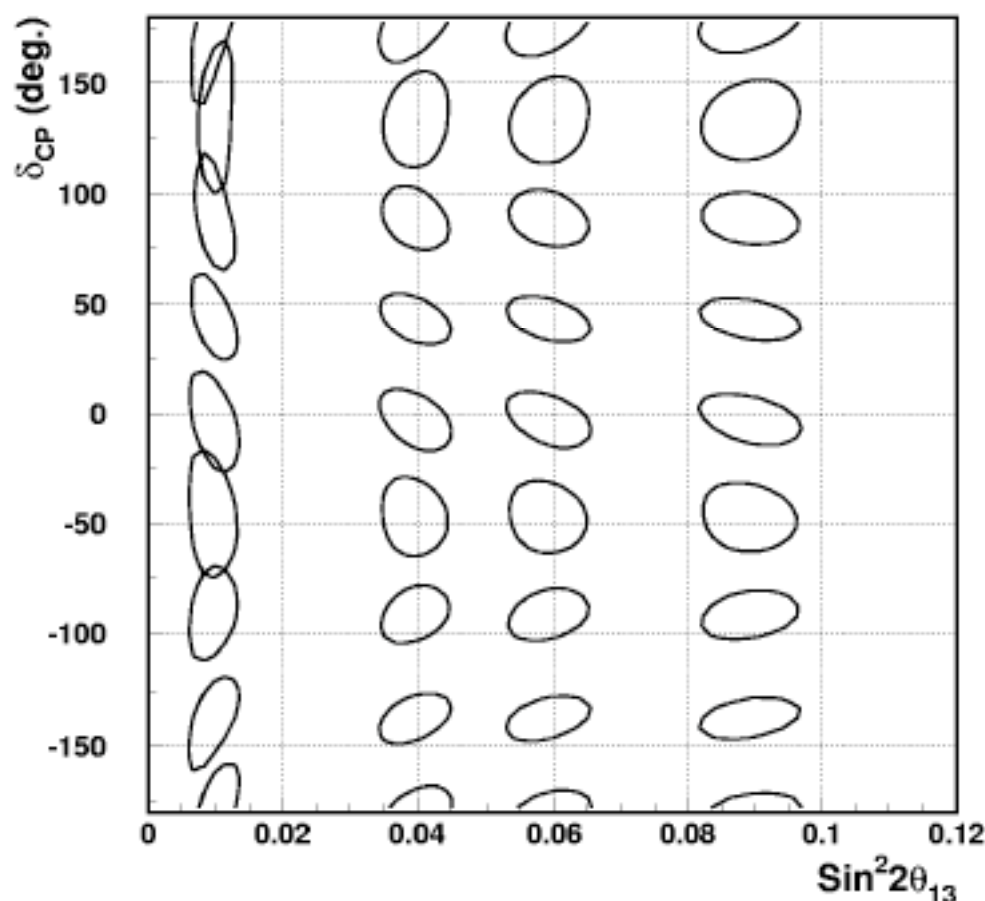


Figure 6: 90% confidence level error contours in $\text{sin}^2 2\theta_{13}$ versus δ_{CP} for statistical and systematic errors with neutrino data alone. Left is for BNL-HS and right is for FNAL-HS. The test point used here is $\text{sin}^2 2\theta_{13} = 0.04$ and $\delta_{CP} = 45^\circ$. $\Delta m^2_{32} = 0.0025 eV^2$, and $\Delta m^2_{21} = 7.3 \times 10^{-5} eV^2$. The values of $\text{sin}^2 2\theta_{12}$ and $\text{sin}^2 2\theta_{23}$ are set to 0.86, 1.0, respectively.

Important Considerations

Regular hierarchy ν and Antiv ν running



If signal is well above background CP resolution is indep. of $\sin^2 2\theta_{13}$

Wide band beam and 2540 km eliminate many parameter correlations.

For 3-generation mixing only neutrino running is needed. Anti-neutrino running gives better precision or New physics.

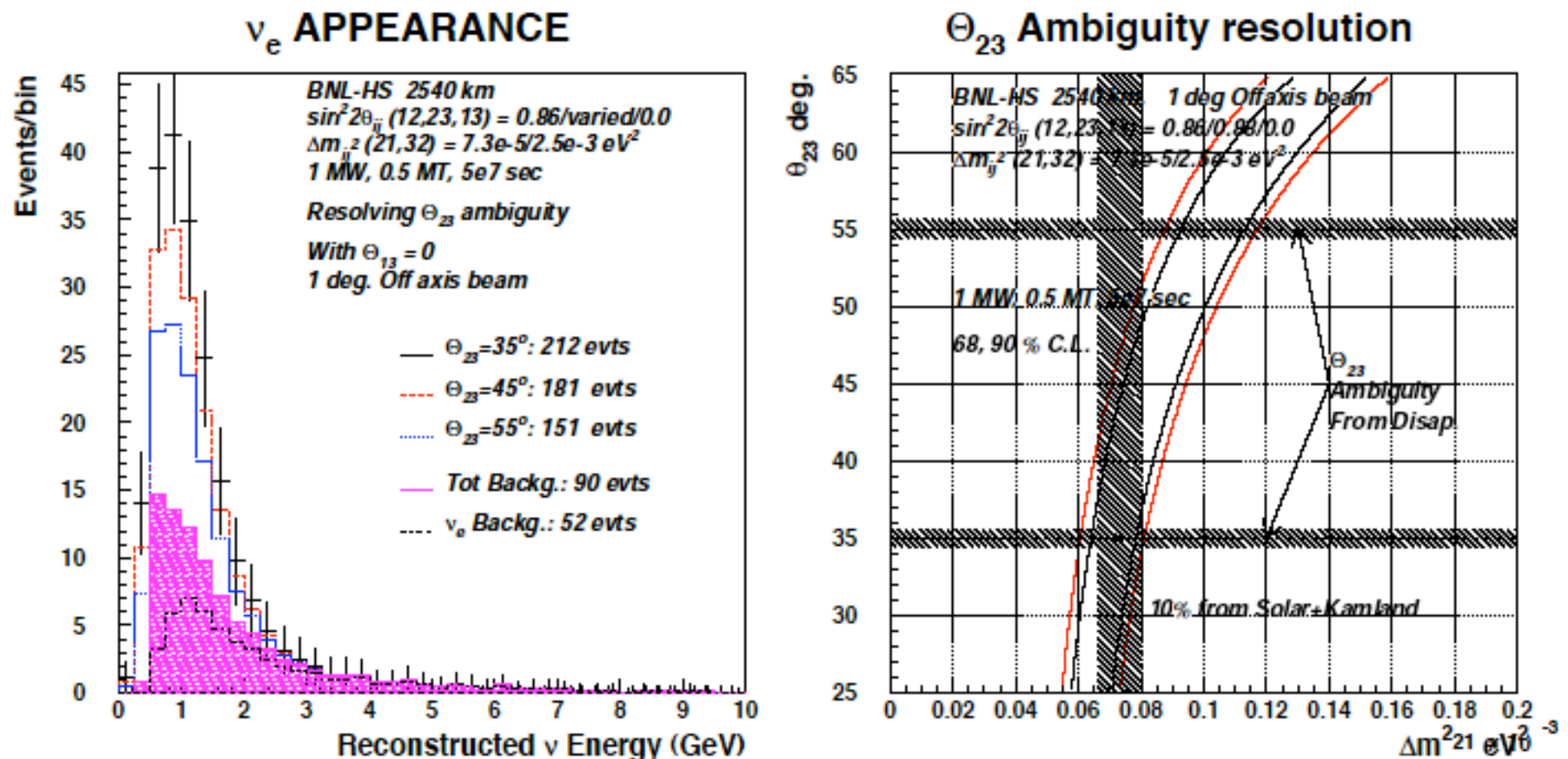


Figure 8: Expected spectrum of electron neutrinos (left) for $\theta_{13} = 0$ and other assumed parameters indicated in the figure. The right hand side shows the resolution of the $\theta_{23} \rightarrow \pi/2 - \theta_{23}$ ambiguity using the measurement of $\sin^2 2\theta_{23}$ from disappearance and assuming a 10% measurement of Δm_{21}^2 from KAMLAND. The area between the curves is allowed by the appearance spectrum (left) for $\theta_{23} = 35^\circ$.

Conclusions

- Neutrino physics entering new phase.
- We can now ask deep questions:
 - Mass: are neutrinos own anti-particles ?
Do neutrinos violate CP conservation ?
Relationship of quarks and neutrinos ?
- New facilities of intense beams and large detectors are needed: APS neutrino study.